## Holographic description of $\mathrm{SCFT}_{5}$ compactifications

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Abstract: We present three infinite families of supersymmetric Type IIB backgrounds with $\mathrm{AdS}_{4}, \mathrm{AdS}_{3}$ and $\mathrm{AdS}_{2}$ factors, dual to SCFTs in 3, 2 and 1 space-time dimensions respectively. These field theories emerge at low energies, after a twisted compactification of a family of five dimensional $\mathcal{N}=1$ SCFTs on hyperbolic spaces. The holographic flows across dimensions are explicitly computed. We also discuss a family of SUSY breaking backgrounds, dual to a QCD-like quiver with massive (bi)fundamental matter. Some field theoretical observables are computed for these theories at the fixed points and along the flow.

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## 1 Introduction

The construction of half-maximal-BPS backgrounds with an AdS-factor for the Type II string or for M-theory is a very fertile problem. Illuminated by Maldacena's AdS/CFT correspondence [1], this problem gains significance as the study of non-perturbative aspects of conformal field theories preserving eight Poincaré SUSYs in diverse dimensions. By now, there is a beautiful correspondence between infinite families of $\operatorname{AdS}_{D} \times S^{2} \times \Sigma_{8-D}$ solutions preserving eight Poincaré SUSYs $(D=2, \ldots, 7)$ and SCFTs in dimension $d=D-1$ with SU(2) R-symmetry.

In fact, for the case $D=2,3,4,5,6,7$ the formalism, backgrounds and dual field theories are respectively described in the papers [2]-[6] (for $\operatorname{AdS}_{2}$ ), [7]-[15] (for $\mathrm{AdS}_{3}$ ), [16][24] (for $\mathrm{AdS}_{4}$ ), [25]-[30] (for $\mathrm{AdS}_{5}$ ), [31]-[42] (for $\mathrm{AdS}_{6}$ ) and [43]-[48] (for $\mathrm{AdS}_{7}$ ).

A natural next step is to achieve the same classification of Type II (or M-theory) backgrounds and CFTs in situations with less SUSY, typically breaking also part of the $\mathrm{SU}(2)$ R-symmetry. Establishing the correspondence between supergravity backgrounds and precise QFT/CFTs (in the less SUSY cases) is a very interesting and demanding problem. Few papers have attempted this, mostly due to the technical difficulties in solving BPS equations (generically, non-linear and coupled PDEs). The achievements are less spectacular than those in the more symmetric circumstances. See [49]-[53], for some works in this direction.

On the field theory side, a popular way of constructing SCFTs is to consider twisted compactifications of the $d=6(0,2)$ SCFT on different manifolds. The $(0,2)$ theory then acts as a 'mother' of the lower dimensional ones (for example, the class $\mathcal{S}$ theories).

In five dimensions there is an infinite family of SCFTs with eight Poincaré SUSYs. Its holographic description is given in the papers [31]-[42]. It is natural to ask if this family of SCFTs can be compactified on two, three, and four manifolds leading to lower dimensional SCFTs. This problem is ideal to tackle using AdS/CFT techniques.

Indeed, a technical tool used in carving the space of possible string backgrounds dual to SCFTs, is to find solutions in D-dimensional AdS-gauged supergravity and lift them to Type II or M-theory. These solutions are typically of the form $\operatorname{AdS}_{D-p} \times \Sigma_{p}$, where $\Sigma_{p}$ is a compact space, and contain a host of scalars and forms excited. Similar procedures are used to construct duals to defect conformal field theories, see for example [54, 55], or massive gravity models [56].

This technique suggests that the dual field theory is a twisted compactification of a ( $D-1$ ) dimensional CFT that reaches a strongly-coupled lower dimensional IR fixed point of dimension $(D-p-1)$. In this paper, we are interested in finding the flow to these IR fixed points dual to lower dimensional SCFTs.

More concretely, we consider five-dimensional $\mathcal{N}=1$ SCFTs (for example, the strongly coupled fixed point of 5 d linear quiver field theories) dual to Type IIB solutions with $\mathrm{AdS}_{6} \times$ $S^{2}$ factors. Holographically, the RG flow to a lower dimensional SCFT is characterized by a solution in an effective six dimensional gauged supergravity [57] that interpolates between $\mathrm{AdS}_{6} \rightarrow \mathrm{AdS}_{d+1} \times \Sigma_{5-d}$. This describes the strongly coupled dynamics of the 5d SCFT compactified on $\Sigma_{5-d}$ that reaches a $d$-dimensional $\mathrm{CFT}_{d}$, with less SUSY than the 5d UV one. In this work, we describe these RG-flows, study fixed points and calculate quantities
characterising both the flows and the fixed points. In the same line, SUSY breaking compactifications can be considered.

The contents of this paper are distributed as follows: in section 2, we summarise the needed aspects of Romans' six dimensional $F_{4}$ gauged supergravity. In section 3 we discuss solutions to this gauged supergravity that preserve some fraction of SUSY and have $\operatorname{AdS}_{4}$, $\mathrm{AdS}_{3}$ and $\mathrm{AdS}_{2}$ spaces. We also study compactifications on a circle with SUSY breaking boundary conditions.

Section 4 carefully describes two formulations of the lift to Type IIB of the solutions presented in section 3. We use both the language of [32-35] and [42] and show the equivalence of the two formulations. In section 5 the three new SUSY preserving families of Type IIB backgrounds with $\mathrm{AdS}_{4}, \mathrm{AdS}_{3}$ and $\mathrm{AdS}_{2}$ factors are explicitly written. These infinite families are labelled by a function $V(\sigma, \eta)$ that also encodes the dual five-dimensional $\mathcal{N}=2$ SCFT undergoing compactification. Some field theoretical aspects of the lower dimensional $\mathrm{SCFT}_{d}$ are discussed in section 6. Among other things the number of degrees of freedom is defined and computed in detail. The holographic central charge is calculated at the fixed points and we also present an analog monotonic observable which characterise the dimensional flows. Interestingly, in terms of the quiver parameters, the holographic central charge of the IR fixed point is proportional to the UV one, confirming the picture advocated in [58]. We also present a dual to a QCD-like quiver theory with massive matter. Section 7 presents a summary and conclusions of this work. Some appendices with technical details complement the presentation.

## 2 Summary of six dimensional Romans $F_{4}$ supergravity

In this section we give an account of six dimensional Romans' $F_{4}$ gauged supergravity [57]. We use slightly different conventions (and a different signature) to those in [57], in order to profit from the lift to Type IIB presented in [59], as we discuss in section 4.

The six dimensional Romans' $\mathrm{F}(4)$ gauged supergravity is defined in terms of a real scalar field $X$, a three-form

$$
\begin{equation*}
F_{3}=d A_{2}, \tag{2.1}
\end{equation*}
$$

an Abelian gauge field $A_{1}$ with field strength

$$
\begin{equation*}
F_{2}=d A_{1}+\frac{2}{3} \tilde{g} A_{2}, \tag{2.2}
\end{equation*}
$$

and a non-Abelian $\operatorname{SU}(2)$ gauge field $A^{i}$ with curvature

$$
\begin{equation*}
F^{i}=d A^{i}+\frac{1}{2} \epsilon^{i j k} A^{j} \wedge A^{k} \tag{2.3}
\end{equation*}
$$

The parameter $\tilde{g}$ is a coupling in the 6 d theory. The units associated to the various fields and curvatures are,

$$
\begin{equation*}
\left[A_{2}\right]=\left[A_{1}\right]=1, \quad\left[F_{2}\right]=\left[F_{3}\right]=m, \quad\left[A^{i}\right]=m,\left[F^{i}\right]=m^{2}, \quad[X]=1, \quad[\tilde{g}]=m \tag{2.4}
\end{equation*}
$$

The bosonic part of the lagrangian reads,

$$
\begin{align*}
\mathcal{L}= & R *_{6} 1+4 \frac{*_{6} d X \wedge d X}{X^{2}}-\tilde{g}^{2}\left(\frac{2}{9} X^{-6}-\frac{8}{3} X^{-2}-2 X^{2}\right) *_{6}  \tag{2.5}\\
& +\frac{1}{2} X^{4} *_{6} F_{3} \wedge F_{3}-\frac{1}{2} X^{-2}\left(*_{6} F_{2} \wedge F_{2}+\frac{1}{\tilde{g}^{2}} *_{6} F^{i} \wedge F^{i}\right) \\
& -\frac{1}{2} \tilde{A}_{2} \wedge\left(d A_{1} \wedge d A_{1}+\frac{2}{3} \tilde{g} d A_{1} \wedge A_{2}+\frac{4}{27} \tilde{g}^{2} A_{2} \wedge A_{2}+\frac{1}{\tilde{g}^{2}} F^{i} \wedge F^{i}\right)
\end{align*}
$$

The equations of motion are

$$
\begin{align*}
d\left(X^{4} *_{6} F_{3}\right)= & \frac{1}{2} F_{2} \wedge F_{2}+\frac{1}{2 \tilde{g}^{2}} F^{i} \wedge F^{i}+\frac{2}{3} \tilde{g} X^{-2} *_{6} F_{2}  \tag{2.6}\\
d\left(X^{-2} *_{6} F_{2}\right)= & -F_{2} \wedge F_{3}  \tag{2.7}\\
D\left(X^{-2} *_{6} F^{i}\right)= & -F_{3} \wedge F^{i}  \tag{2.8}\\
d\left(X^{-1} *_{6} d X\right)= & \frac{1}{4} X^{4} *_{6} F_{3} \wedge F_{3}-\frac{X^{-2}}{8}\left(*_{6} F_{2} \wedge F_{2}+\frac{1}{\tilde{g}^{2}} *_{6} F^{i} \wedge F^{i}\right) \\
& -\tilde{g}^{2}\left(\frac{1}{6} X^{-6}-\frac{2}{3} X^{-2}+\frac{1}{2} X^{2}\right) *_{6} 1 \tag{2.9}
\end{align*}
$$

where we defined $D$ as the $\mathrm{SU}(2)$ gauge-covariant derivative, defined on a differential for $C^{i}$ by

$$
\begin{equation*}
D C^{i}=d C^{i}+\epsilon_{i j k} A^{j} \wedge C^{k} \tag{2.10}
\end{equation*}
$$

Finally, the Einstein's equations are

$$
\begin{align*}
R_{\mu \nu}= & 4 X^{-2} \partial_{\mu} X \partial_{\nu} X+\tilde{g}^{2}\left(\frac{1}{18} X^{-6}-\frac{2}{3} X^{-2}-\frac{1}{2} X^{2}\right) g_{\mu \nu}+\frac{X^{4}}{4}\left(F_{3 \mu} \cdot F_{3 \nu}-\frac{1}{6} g_{\mu \nu} F_{3}^{2}\right) \\
& +\frac{X^{-2}}{2}\left(F_{2 \mu} \cdot F_{2 \nu}-\frac{1}{8} g_{\mu \nu} F_{2}^{2}\right)+\frac{X^{-2}}{2 \tilde{g}^{2}}\left(F_{2 \mu}^{i} \cdot F_{2 \nu}^{i}-\frac{1}{8} g_{\mu \nu}\left(F_{2}^{i}\right)^{2}\right) \tag{2.11}
\end{align*}
$$

where $F_{\mu}=\iota_{\mu} F$ is the contraction along the direction $\partial_{\mu}, F \cdot G=F_{\mu_{1} \ldots \mu_{p}} G^{\mu_{1} \ldots \mu_{p}}$, and $F^{2}=F \cdot F$.

The fermionic part of the Lagrangian and SUSY transformations can be found in [57]. The simplest possible solution sets all matter fields to zero $A_{2}=A_{1}=A^{i}=0$ and $X=1$. The metric is that of $\mathrm{AdS}_{6}$ of radius $R^{2}=\frac{2}{9 \tilde{g}^{2}}$. This solution preserves eight Poincaré supercharges. Let us now discuss various solutions of this six dimensional system.

## 3 Solutions in six dimensional $F_{4}$ supergravity

In this section we discuss various solutions to $F_{4}$ gauged supergravity. The fixed point solutions described below are not new, in fact they can be found in [60] and [61]. We present them in detail in our conventions and also display new numerical interpolations between fixed points. In all cases we have checked that all the equations (2.6)-(2.11) are satisfied.

These solutions describe spontaneous compactifications, triggered by fluxes, of the $\mathrm{AdS}_{6}$ SUSY vacuum of six dimensional supergravity on the manifolds $H_{2}, H_{3}, H_{2} \times H_{2}$. They are discussed in sections $3.1,3.2,3.3$ respectively. These solutions preserve different


Figure 1. Numerical interpolation between the $A d S_{6}$ and $A d S_{4}$ fixed point. The details of the numerical analysis are explained in appendix B.1. We do not plot $g(r)$ since it is given by $g(r)=$ $-f(r)+\log X(r)$.
fractions of the original Poincaré SUSYs. Finally, a non-SUSY solution obtained as a double Wick rotation of an $\mathrm{AdS}_{6}$ black hole is described in section 3.4.

We start by writing flows between $\mathrm{AdS}_{6}$ and $A d S_{4}$ when the supergravity compactifies on a compact hyperbolic two-dimensional space

## 3.1 $\quad \mathrm{AdS}_{6} \rightarrow \mathrm{AdS}_{4} \times \mathrm{H}_{2}$

This configuration was first discussed in [60], we review it here in our conventions. We propose a metric

$$
\begin{equation*}
d s^{2}=e^{2 f(r)}\left(-d t^{2}+d y_{1}^{2}+d y_{2}^{2}+e^{2 g(r)} d r^{2}\right)+e^{2 h(r)} \frac{\left(d z^{2}+d x^{2}\right)}{z^{2}} \tag{3.1}
\end{equation*}
$$

where $g(r)$ is an arbitrary function which takes into account the reparameterisation invariance of the coordinate $r$. The gauge field configuration breaks $\mathrm{SU}(2)$ to $\mathrm{U}(1)$

$$
\begin{equation*}
A^{1}=0, \quad A^{2}=0, \quad A^{3}=\frac{1}{z} d x \tag{3.2}
\end{equation*}
$$

The dilaton depends only on the $r$ coordinate, $X=X(r)$ and the fields $A_{1}=A_{2}=0$. The BPS and the equations of motion are solved if we impose

$$
\begin{align*}
X^{\prime} & =-\frac{e^{f+g} \tilde{g}}{2 \sqrt{2}}\left[X^{-2}-X^{2}-\frac{1}{2 \tilde{g}^{2}} e^{-2 h}\right]  \tag{3.3}\\
h^{\prime} & =-\frac{e^{f+g} \tilde{g}}{2 \sqrt{2} X}\left[\frac{X^{-2}}{3}+X^{2}-\frac{3}{2 \tilde{g}^{2}} e^{-2 h}\right]  \tag{3.4}\\
f^{\prime} & =-\frac{e^{f+g} \tilde{g}}{2 \sqrt{2} X}\left[\frac{X^{-2}}{3}+X^{2}+\frac{1}{2 \tilde{g}^{2}} e^{-2 h}\right] \tag{3.5}
\end{align*}
$$

The solutions to these equations generate backgrounds preserving four Poincaré SUSYs $(\mathcal{N}=2$ in the language of the dual 3d SCFT), see [60] for a detailed analysis. The BPS
equations admit a fixed point solution by requiring that both $X$ and $h$ are constants,

$$
\begin{equation*}
X^{4}=\frac{2}{3}, \quad e^{2 h}=\sqrt{\frac{3}{2}} \tilde{g}^{-2}, \quad e^{2 g}=e^{-2 f} \frac{\sqrt{6}}{\tilde{g}^{2}}\left(f^{\prime}\right)^{2} . \tag{3.6}
\end{equation*}
$$

Plugging this result in eq. (3.1), the metric is $\operatorname{AdS}_{4} \times H_{2}$ (where we can now use $f$ as a coordinate). Also notice that for small $r$, the expansion

$$
\begin{equation*}
f=-\log r+O\left(r^{2}\right), \quad h=-\log r+O\left(r^{2}\right), \quad X=1+O\left(r^{2}\right), \quad g=O\left(r^{2}\right) \tag{3.7}
\end{equation*}
$$

gives a power series solution to the BPS equations which approaches $\operatorname{AdS}_{6}$ as $r \rightarrow 0 .{ }^{1}$
The solution which interpolates among these two fixed points is shown in figure 1. For the numerical analysis we have set $\tilde{g}=3 / \sqrt{2}$ and $g(r)=-f(r)+\log X(r)$. With this choice of radial coordinate - different from the one in eq. (3.7), we find the $\mathrm{AdS}_{4}$ fixed point at $r \rightarrow \infty$ and the $\mathrm{AdS}_{6}$ one at $r \rightarrow-\infty$. Indeed, at the $\mathrm{AdS}_{4}$ fixed point we have

$$
\begin{equation*}
r \gg 0 \quad X=\sqrt[4]{\frac{2}{3}}=0.90, \quad h=\frac{1}{4} \log \frac{2}{3^{3}}=-0.65, \quad f=-\sqrt{\frac{3}{2}} r \tag{3.8}
\end{equation*}
$$

consistently with eq. (3.6). In the $\mathrm{AdS}_{6}$ fixed point we have

$$
\begin{equation*}
r \ll 0 \quad X=1, \quad h=-r, \quad f=-r \tag{3.9}
\end{equation*}
$$

which for $r \rightarrow-\infty$ approaches the metric of $\mathrm{AdS}_{6}$. See figure 1 for a plot of the numerical solutions. Details of the numerical analysis can be found in appendix B.1. Let us now discuss spontaneous compactifications to $\mathrm{AdS}_{3}$.

## $3.2 \quad \mathrm{AdS}_{6} \rightarrow \mathrm{AdS}_{3} \times \mathrm{H}_{3}$

In this section we discuss a flow from $\mathrm{AdS}_{6}$ to a solution of the form $\mathrm{AdS}_{3} \times H_{3}$, first found in [60]. Interestingly, this background needs the non-Abelian character of the gauge field $A^{i}$. The six dimensional configuration reads

$$
\begin{align*}
d s^{2} & =e^{2 f(r)}\left(-d t^{2}+d y^{2}+e^{2 g(r)} d r^{2}\right)+e^{2 h(r)} \frac{\left(d x_{1}^{2}+d x_{2}^{2}+d z^{2}\right)}{z^{2}},  \tag{3.10}\\
A^{1} & =-\frac{1}{z} d x_{1}, \quad A^{2}=0, \quad A^{3}=-\frac{1}{z} d x_{2}, \quad X=X(r) .
\end{align*}
$$

All the other fields are set to zero. The BPS equations are

$$
\begin{align*}
X^{\prime} & =\frac{\tilde{g} e^{f+g}}{4 \sqrt{2}}\left(2 X^{2}-2 X^{-2}+\frac{3}{\tilde{g}^{2}} e^{-2 h}\right),  \tag{3.11}\\
f^{\prime} & =-\frac{\tilde{g} e^{f+g}}{4 \sqrt{2} X}\left(2 X^{2}+\frac{2 X^{-2}}{3}+\frac{3}{\tilde{g}^{2}} e^{-2 h}\right),  \tag{3.12}\\
h^{\prime} & =-\frac{\tilde{g} e^{f+g}}{4 \sqrt{2} X}\left(2 X^{2}+\frac{2 X^{-2}}{3}-\frac{5}{\tilde{g}^{2}} e^{-2 h}\right) . \tag{3.13}
\end{align*}
$$

Any solution of the BPS equations (3.11)-(3.13) is a solution of the equations of motion (2.6)-(2.11). These backgrounds will preserve two Poincaré SUSYs, as analysed in [60].

[^0]

Figure 2. Numerical interpolation between the $A d S_{6}$ and $A d S_{3}$ fixed point. The details of the numerical analysis are explained in appendix B.2. Again, $g$ is given by $g=-f+\log X$.

One possible solution is $X=1+O\left(r^{2}\right)$ for small $r$ and $f \sim h \sim-\log r$. This reproduces $\mathrm{AdS}_{6}$ (for $r \rightarrow 0$ ). There is also a fixed point solution to the equations,

$$
\begin{equation*}
X^{4}=\frac{1}{2}, \quad e^{2 h}=\frac{3}{\sqrt{2} \tilde{g}^{2}}, \quad e^{2 g}=e^{-2 f} \frac{9}{2^{5 / 2}} \frac{1}{\tilde{g}^{2}}\left(f^{\prime}\right)^{2} \tag{3.14}
\end{equation*}
$$

leading to a space of the form $\mathrm{AdS}_{3} \times H_{3}$, with a fixed size $H_{3}$.
The solution describing the flow $\mathrm{AdS}_{6} \rightarrow \mathrm{AdS}_{3} \times H_{3}$ is plotted in figure 2. The details of the numerical analysis are written in appendix B.2. As above, we set $\tilde{g}=3 / \sqrt{2}$ and we use a convenient radial coordinate set by the choice $g(r)=-f(r)+\log X(r)$. Using this radial coordinate the $\mathrm{AdS}_{6}$ fixed point is found as $r \rightarrow-\infty$, while the $\mathrm{AdS}_{3}$ fixed point is found at $r \rightarrow+\infty$. Indeed, we have

$$
\begin{equation*}
r \gg 0 \quad X=\sqrt[4]{\frac{1}{2}}=0.84, \quad h=\frac{1}{4} \log \frac{2}{9}=-0.38, \quad f=-\sqrt{2} r \tag{3.15}
\end{equation*}
$$

which is consistent with eq. (3.14). In the $\mathrm{AdS}_{6}$ fixed point we have

$$
\begin{equation*}
r \ll 0 \quad X=1, \quad h=-r, \quad f=-r \tag{3.16}
\end{equation*}
$$

Let us now move to analyse spontaneous compactifications to $\mathrm{AdS}_{2}$.

## $3.3 \quad \mathrm{AdS}_{6} \rightarrow \mathrm{AdS}_{2} \times H_{2}^{(1)} \times H_{2}^{(2)}$

The backgrounds we summarise in this section were first discussed in [61]. As the solution in section 3.1 , he configuration breaks the $\mathrm{SU}(2)$ gauge symmetry to $\mathrm{U}(1)$, but in this case we also have $A_{2}$ turned on

$$
\begin{align*}
d s_{6}^{2} & =e^{2 f(r)}\left(-d t^{2}+e^{2 g(r)} d r^{2}\right)+e^{2 h_{1}(r)} \frac{\left(d z_{1}^{2}+d x_{1}^{2}\right)}{z_{1}^{2}}+e^{2 h_{2}(r)} \frac{\left(d z_{2}^{2}+d x_{2}^{2}\right)}{z_{2}^{2}}  \tag{3.17}\\
A^{1} & =0, \quad A^{2}=0, \quad A^{3}=\frac{1}{z_{1}} d x_{1}+\frac{1}{z_{2}} d x_{2}, \quad X=X(r)  \tag{3.18}\\
A_{3} & =0, \quad A_{2}=\frac{9}{4 \tilde{g}^{4}} X(r)^{2} e^{2\left(f-h_{1}-h_{2}\right)+g} d t \wedge d r \tag{3.19}
\end{align*}
$$



Figure 3. Numerical interpolation between the $A d S_{6}$ and $A d S_{2}$ fixed point. The details of the numerical analysis are explained in appendix B.3. Again, $g$ is given by $g=-f+\log X$ and we have assumed $h_{1}=h_{2}$ along the flow.

With these definitions and using eqs. (2.1)-(2.2), we have

$$
\begin{equation*}
F_{3}=d A_{2}=0, \quad F_{2}=\frac{2}{3} \tilde{g} A_{2} \neq 0 \tag{3.20}
\end{equation*}
$$

The amount of SUSY preserved and the BPS system for this configuration were studied in [61]. The BPS equations read,

$$
\begin{align*}
h_{1}^{\prime} & =-\frac{\tilde{g} e^{f+g}}{2 \sqrt{2} X}\left(X^{2}+\frac{X^{-2}}{3}-\frac{1}{2 \tilde{g}^{2}}\left(3 e^{-2 h_{1}}-e^{-2 h_{2}}\right)-\frac{3 X^{2}}{4 \tilde{g}^{4}} e^{-2 h_{1}-2 h_{2}}\right)  \tag{3.21}\\
h_{2}^{\prime} & =-\frac{\tilde{g} e^{f+g}}{2 \sqrt{2} X}\left(X^{2}+\frac{X^{-2}}{3}-\frac{1}{2 \tilde{g}^{2}}\left(3 e^{-2 h_{2}}-e^{-2 h_{1}}\right)-\frac{3 X^{2}}{4 \tilde{g}^{4}} e^{-2 h_{1}-2 h_{2}}\right)  \tag{3.22}\\
f^{\prime} & =-\frac{\tilde{g} e^{f+g}}{2 \sqrt{2} X}\left(X^{2}+\frac{X^{-2}}{3}+\frac{1}{2 \tilde{g}^{2}}\left(e^{-2 h_{1}}+e^{-2 h_{2}}\right)+\frac{9 X^{2}}{4 \tilde{g}^{4}} e^{-2 h_{1}-2 h_{2}}\right)  \tag{3.23}\\
X^{\prime} & =-\frac{\tilde{g} e^{f+g}}{2 \sqrt{2}}\left(-X^{2}+X^{-2}-\frac{1}{2 \tilde{g}^{2}}\left(e^{-2 h_{1}}+e^{-2 h_{2}}\right)+\frac{3 X^{2}}{4 \tilde{g}^{4}} e^{-2 h_{1}-2 h_{2}}\right) \tag{3.24}
\end{align*}
$$

The system admits a fixed-point solution

$$
\begin{equation*}
e^{h_{1}}=e^{h_{2}}=\sqrt[4]{\frac{3}{2}} \tilde{g}^{-1}, \quad X=\sqrt[4]{\frac{2}{3}}, \quad e^{2 f+2 g}=\sqrt{\frac{3}{2}} \frac{\left(f^{\prime}\right)^{2}}{2 \tilde{g}^{2}} \tag{3.25}
\end{equation*}
$$

which correspond to an $\mathrm{AdS}_{2} \times H_{2} \times H_{2}$ metric. The BPS system can also be solved with power series expansion around $r \sim 0$ whose leading terms are

$$
\begin{equation*}
f(r)=-\log r, \quad h_{1}(r)=-\log r, \quad h_{2}(r)=-\log r, \quad X(r)=1, \quad g(r)=0 \tag{3.26}
\end{equation*}
$$

which approaches $\mathrm{AdS}_{6}$ as $r \rightarrow 0$.
Since $h_{1}=h_{2}$ at both fixed points, we apply the simplifying ansatz $h_{1}=h_{2}$ along the full $\mathrm{AdS}_{6} \rightarrow \mathrm{AdS}_{2} \times H_{2} \times H_{2}$ flow. Notice that this choice makes one equation in the BPS
system redundant. The flow is plotted in figure 3 and we refer to appendix B. 3 for the details about the numerical analysis. Again we set $\tilde{g}=3 / \sqrt{2}$ and $g(r)=-f(r)+\log X(r)$; the $\mathrm{AdS}_{2}$ fixed point is found at $r \rightarrow+\infty$ :

$$
\begin{equation*}
r \gg 0 \quad X=\sqrt[4]{\frac{2}{3}}=0.90, \quad h_{i}=\frac{1}{4} \log \frac{2}{27}=-0.65, \quad f=-\sqrt{6} r \tag{3.27}
\end{equation*}
$$

which is consistent with eq. (3.25), while we obtain the $\mathrm{AdS}_{6}$ for

$$
\begin{equation*}
r \ll 0 \quad X=1, \quad h_{i}=-r, \quad f=-r . \tag{3.28}
\end{equation*}
$$

Finally, let us discuss a solution describing the compactification on a circle.

## $3.4 \quad \mathrm{AdS}_{6} \rightarrow R^{1,3} \times S^{1}$

In this section we discuss a very simple solution. It can be obtained by Wick rotating a non supersymmetric black hole in $\mathrm{AdS}_{6}$. In fact, consider the solution with all the fields turned off $A_{2}=A_{1}=A^{i}=0$ and $X=1$. The metric is,

$$
\begin{equation*}
d s^{2}=e^{2 \rho}\left(-d \tau^{2}+d \vec{x}_{3}^{2}\right)+\frac{d \rho^{2}}{g(\rho)}+e^{2 \rho} g(\rho) d \psi^{2} . \tag{3.29}
\end{equation*}
$$

If $g(\rho)=1$ this is the supersymmetric $\mathrm{AdS}_{6}$ solution mentioned below eq. (2.11). Here we set

$$
\begin{equation*}
g(\rho)=1-e^{5\left(\rho_{*}-\rho\right)} . \tag{3.30}
\end{equation*}
$$

The equations of motion (2.6)-(2.11) are satisfied. Supersymmetry is broken, unless $\rho_{*} \rightarrow$ $-\infty$. The variable $\rho$, ranges in $\left[\rho_{*}, \infty\right)$. For this background to be smooth we need to set up constants such that a conical singularity is avoided. This fixes $\rho_{*}=\log \left(\frac{2}{5 \sqrt{5}}\right)$. In fact, close to the end of the space, $x=\left(\rho-\rho_{*}\right) \rightarrow 0$,

$$
\frac{d \rho^{2}}{g(\rho)}+e^{2 \rho} g(\rho) d \psi^{2} \approx \frac{d x^{2}}{5 x}+5 x e^{2 \rho_{*}} d \psi^{2}=d u^{2}+\frac{125 e^{2 \rho_{*}}}{4} u^{2} d \psi^{2}=d u^{2}+u^{2} d \psi^{2} .
$$

In the next section, we consider all the six-dimensional backgrounds presented in this section and lift them to Type IIB. We will use the lifting formulas of [59], written in the formulation of $[42,62]$. This will provide new supersymmetric Type IIB families of backgrounds with $\mathrm{AdS}_{4}, \mathrm{AdS}_{3}$, and $\mathrm{AdS}_{2}$ factors. After lifting, the solution in section 3.4 provides a non-SUSY background interpolating between $\operatorname{AdS}_{6}$ and $R^{1,3} \times S^{1}$ in Type IIB. After that, in section 6, we present a dual CFT/QFT interpretation of these backgrounds and compute interesting observables.

Before all of this, we carefully describe the uplifting formulas and the dictionary between the formalism of [59] and that of [42, 62]. To this we turn.

## 4 Type IIB $\mathrm{AdS}_{6}$ background

In this section we present two different but equivalent ways of writing an infinite family of $\mathrm{AdS}_{6}$ background in type IIB supergravity. These backgrounds preserve eight Poincaré
supercharges. Aside from the $\mathrm{SO}(2,5)$ isometry realised by the $\mathrm{AdS}_{6}$ factor, the $\mathrm{SU}(2)$ isometries of a two-sphere are also imposed. The internal space has just two unconstrained dimensions.

It is convenient to parameterise the two-dimensional space using complex coordinates. This leads to expressing the background in terms of two holomorphic functions. This infinite family of solutions was first written in [31]. In this paper, we use the notation of [59].

A second possibility, first analysed in [62] and further developed in [42], is to express the background in terms of a real function which solves a Laplace equation. Bellow we describe both formulations and show their equivalence.

### 4.1 Using holomorphic functions

This first formulation applies in the case for which all the fields in the background are locally written in terms of two unconstrained holomorphic functions $\mathcal{A}_{ \pm}(z)$ of a complex coordinate $z$. To succinctly describe the formalism is convenient to define:

- An auxiliary holomorphic function $\mathcal{B}$ in terms of the differential equation,

$$
\begin{equation*}
\partial_{z} \mathcal{B}=\mathcal{A}_{+} \partial_{z} \mathcal{A}_{-}-\mathcal{A}_{-} \partial_{z} \mathcal{A}_{+} . \tag{4.1}
\end{equation*}
$$

- The real functions,

$$
\begin{array}{ll}
\mathcal{G}=\left|\mathcal{A}_{+}\right|^{2}-\left|\mathcal{A}_{-}\right|^{2}+2 \operatorname{Re} \mathcal{B}, & \kappa^{2}=-\partial_{z} \partial_{\bar{z}} \mathcal{G}=\left|\partial_{z} \mathcal{A}_{-}\right|^{2}-\left|\partial_{z} \mathcal{A}_{+}\right|^{2}, \\
\mathcal{Y}=\frac{\kappa^{2} \mathcal{G}}{\left|\partial_{z} \mathcal{G}\right|^{2}}, & \mathcal{D}=1+\frac{2}{3 \mathcal{Y}} . \tag{4.2}
\end{array}
$$

Using this, the string-frame metric reads

$$
\begin{equation*}
d s_{s t}^{2}=e^{\frac{\Phi}{2}}\left(f_{1} d s^{2}\left(\operatorname{AdS}_{6}\right)+f_{2} d s^{2}\left(S^{2}\right)+4 f_{3} d z d \bar{z}\right), \tag{4.3}
\end{equation*}
$$

where

$$
\begin{equation*}
f_{1}=\frac{\kappa^{2} \sqrt{\mathcal{D}}}{f_{3}}, \quad f_{2}=\frac{1}{9 f_{3}} \frac{\kappa^{2}}{\sqrt{\mathcal{D}}}, \quad f_{3}^{2}=\frac{\kappa^{4} \sqrt{\mathcal{D}}}{6 \mathcal{G}} . \tag{4.4}
\end{equation*}
$$

The rest of the Type IIB fields, are expressed using the $\operatorname{SU}(1,1)$ covariant formalism,

$$
\begin{align*}
\frac{1+i \tau}{1-i \tau} & =\frac{\mathcal{A}_{+}-\overline{\mathcal{A}}_{-}-\mathcal{C} / \sqrt{\mathcal{D}}}{\overline{\mathcal{A}}_{+}-\mathcal{A}_{-}+\overline{\mathcal{C}} / \sqrt{\mathcal{D}}}, \quad \tau=C_{0}+i e^{-\Phi}, \\
B_{2}+i C_{2} & =\frac{2 i}{9}\left[\frac{\mathcal{C}}{\mathcal{D}}-3\left(\mathcal{A}_{+}+\mathcal{A}_{-}\right)\right] \operatorname{Vol}\left(S^{2}\right), \\
F_{5} & =0 . \tag{4.5}
\end{align*}
$$

We have also defined

$$
\begin{equation*}
\mathcal{C}=\frac{\partial_{z} \mathcal{A}_{+} \partial_{\bar{z}} \mathcal{G}+\partial_{\bar{z}} \overline{\mathcal{A}}_{-} \partial_{z} \mathcal{G}}{\kappa^{2}} . \tag{4.6}
\end{equation*}
$$

In the expressions above, $C_{0}$ and $C_{2}$ are the RR potential for $F_{1}$ and $F_{3}$ respectively while $B_{2}$ is the potential for the NSNS three form $H$. Respect to [59], we have set $c_{6}=1$. In summary, the backgrounds of eqs. (4.1)-(4.6) represent an infinite family of $\mathrm{AdS}_{6} \times S^{2}$ solutions of type IIB preserving eight Poincaré SUSYs. $\mathcal{A}_{ \pm}(z)$.

Next, we present the formulation in terms of one real function [42, 62], and prove the equivalence of this second formulation with the one of [31], above described.

### 4.2 Using a real potential

Another formulation of $\mathrm{AdS}_{6}$ backgrounds with $\mathrm{SU}(2)$ isometry, preserving eight supercharges was given in [62]. All the fields depend on a real potential $V$ which has support on the two-dimensional internal space. We will sketch the relation between this infinite family of backgrounds and the one summarised above in the next section.

It is convenient to parameterise the Riemann surface in terms of two real coordinates, $(\sigma, \eta)$, and use a 'potential function' $V(\sigma, \eta)$ that solves the Laplace partial differential equation

$$
\begin{equation*}
\partial_{\sigma}\left(\sigma^{2} \partial_{\sigma} V\right)+\sigma^{2} \partial_{\eta}^{2} V=0 \tag{4.7}
\end{equation*}
$$

The type IIB background metric in string frame is

$$
\begin{align*}
d s_{10, s t}^{2} & =\tilde{f}_{1}\left[d s^{2}\left(\operatorname{AdS}_{6}\right)+\tilde{f}_{2} d s^{2}\left(S^{2}\right)+\tilde{f}_{3}\left(d \sigma^{2}+d \eta^{2}\right)\right], \quad \tilde{f}_{1}=\frac{2}{3} \sqrt{\sigma^{2}+\frac{3 \sigma \partial_{\sigma} V}{\partial_{\eta}^{2} V}} \\
\tilde{f}_{2} & =\frac{\partial_{\sigma} V \partial_{\eta}^{2} V}{3 \tilde{\Lambda}}, \quad \tilde{f}_{3}=\frac{\partial_{\eta}^{2} V}{3 \sigma \partial_{\sigma} V}, \quad \tilde{\Lambda}=3 \partial_{\eta}^{2} V \partial_{\sigma} V+\sigma\left[\left(\partial_{\eta \sigma}^{2} V\right)^{2}+\left(\partial_{\eta}^{2} V\right)^{2}\right] \tag{4.8}
\end{align*}
$$

The fluxes are given by the following expressions,

$$
\begin{align*}
B_{2} & =\tilde{f}_{4} \operatorname{Vol}\left(S^{2}\right)=\frac{2}{9}\left(\eta-\frac{\left(\sigma \partial_{\sigma} V\right)\left(\partial_{\sigma} \partial_{\eta} V\right)}{\tilde{\Lambda}}\right) \operatorname{Vol}\left(S^{2}\right), \\
C_{2} & =\tilde{f}_{5} \operatorname{Vol}\left(S^{2}\right)=4\left(V-\frac{\sigma \partial_{\sigma} V}{\tilde{\Lambda}}\left(\partial_{\eta} V\left(\partial_{\sigma} \partial_{\eta} V\right)-3\left(\partial_{\eta}^{2} V\right)\left(\partial_{\sigma} V\right)\right)\right) \operatorname{Vol}\left(S^{2}\right), \\
e^{-2 \Phi} & =\tilde{f}_{6}=18^{2} \frac{3 \sigma^{2} \partial_{\sigma} V \partial_{\eta}^{2} V}{\left(3 \partial_{\sigma} V+\sigma \partial_{\eta}^{2} V\right)^{2}} \tilde{\Lambda}, \quad C_{0}=\tilde{f}_{7}=18\left(\partial_{\eta} V+\frac{\left(3 \sigma \partial_{\sigma} V\right)\left(\partial_{\sigma} \partial_{\eta} V\right)}{3 \partial_{\sigma} V+\sigma \partial_{\eta}^{2} V}\right), \\
F_{5} & =0 . \tag{4.9}
\end{align*}
$$

It was shown in [42] that, subject to eq. (4.7), all the equations of motion of the configuration in eqs. (4.8)-(4.9) are satisfied, while supersymmetry is discussed in [62]. Let us now explain the map between these two infinite family of solutions.

### 4.3 Matching the backgrounds

As explained in [42], the two families of solutions discussed in the sections 4.1 and 4.2 actually describe the same configuration. In fact, we can define a complex coordinate from the two real ones $(\sigma, \eta)$

$$
\begin{equation*}
z=\sigma-i \eta \tag{4.10}
\end{equation*}
$$

In terms of this complex coordinate, we choose

$$
\begin{equation*}
\mathcal{A}_{ \pm}=\mp \frac{z}{6}-6 \partial_{z}(\sigma V) \tag{4.11}
\end{equation*}
$$

and we recover the family in eqs. (4.8)-(4.9) from the backgrounds in eqs. (4.1)-(4.6).
Let us further discuss the identification in eq. (4.11). Notice that defining $V=\widehat{V} / \sigma$, eq. (4.7) becomes

$$
\begin{equation*}
\left(\partial_{\sigma}^{2}+\partial_{\eta}^{2}\right) \widehat{V}=0 \tag{4.12}
\end{equation*}
$$

Therefore $\widehat{V}$ is a real harmonic function. Hence it defines one holomorphic function $\mathcal{V}(z)$

$$
\begin{equation*}
\widehat{V}=\mathcal{V}(z)+\overline{\mathcal{V}(z)} . \tag{4.13}
\end{equation*}
$$

It might seem that the backgrounds in section 4.2 are less general than those in section 4.1. Whilst the backgrounds in eqs. (4.1)-(4.6) needed two holomorphic functions $\mathcal{A}_{ \pm}(z)$, those in eqs. (4.8)-(4.9) can be defined with just one holomorphic function $\mathcal{V}(z)$. But, notice that the backgrounds of section 4.1 are invariant under conformal transformations $z \mapsto \mathcal{F}(z)$, which means that one of the holomorphic functions $\mathcal{A}_{ \pm}$can be gauged away.

We can use eq. (4.11) to make more explicit the dictionary between sections 4.1 and 4.2. We have

$$
\begin{align*}
& \mathcal{G}=4 \sigma^{2} \partial_{\sigma} V, \quad \kappa^{2}=2 \sigma \partial_{\eta}^{2} V, \quad \mathcal{Y}=\frac{2 \partial_{\sigma} V \partial_{\eta}^{2} V}{\tilde{\Lambda}-3 \partial_{\eta}^{2} V \partial_{\sigma} V}, \\
& \mathcal{C}=-\frac{\sigma \partial_{\eta \sigma}^{2} V\left(18 \partial_{\eta \sigma}^{2}(\sigma V)-i\right)}{\partial_{\eta}^{2} V}-6 \sigma^{2} \partial_{\eta}^{2} V \tag{4.14}
\end{align*}
$$

Equations (4.11) and (4.14) can be used to write the uplift of a generic solution in Romans' supergravity to type IIB in terms of the potential $V$. Below, we explain this.

## 5 The new families of solutions

In this section we show how to write the uplift of any configuration of Romans supergravity to type IIB. Then, we explicitly write three new infinite families of SUSY type IIB backgrounds with $\mathrm{AdS}_{4}, \mathrm{AdS}_{3}$ and $\mathrm{AdS}_{2}$ factors. We finally discuss the SUSY breaking compactification $\mathrm{AdS}_{6} \rightarrow R^{1,3} \times S^{1}$.

The lift to type IIB of generic configurations of Romans' supergravity was first derived in [59]. ${ }^{2}$ We use eqs. (4.11) and (4.14) to rewrite the result in [59] in terms of the potential $V(\sigma, \eta)$. Whilst we rely on [59] for the general uplift, we have explicitly checked that the type IIB equation of motions are solved for all the cases presented in section 3.

A generic solution in Romans' $F_{4}$ gauged supergravity lifts to a configuration in type IIB given by,

$$
\begin{align*}
d s_{s t}^{2} & =f_{1}\left(d s_{6}^{2}+f_{2} d s^{2}\left(\tilde{S}^{2}\right)+f_{3} d s^{2}\left(R^{2}\right)\right)  \tag{5.1}\\
C_{0} & =f_{7}, \quad e^{-2 \Phi}=f_{6}, \quad F_{5}=4\left(G_{5}+*_{10} G_{5}\right), \\
B_{2} & =f_{4} \operatorname{Vol}\left(\tilde{S}^{2}\right)-\frac{2 \tilde{g}}{9} \sigma F_{2}-\frac{2}{9} \eta y^{i} F^{i}, \\
C_{2} & =f_{5} \operatorname{Vol}\left(\tilde{S}^{2}\right)-4 \tilde{g} \sigma \partial_{\eta} V F_{2}-4 \partial_{\sigma}(\sigma V) y^{i} F^{i},
\end{align*}
$$

where $d s_{6}^{2}$ is defined in terms of the gauged-supergravity metric as

$$
\begin{equation*}
d s_{6}^{2}=\frac{2 \tilde{g}^{2}}{9} d s_{\text {gauged sugra. }}^{2} \tag{5.2}
\end{equation*}
$$

[^1]The functions $f_{i}$ are a deformation of the $\tilde{f}_{i}$ defined in eqs. (4.8)-(4.9):

$$
\begin{array}{ll}
f_{1}=\frac{2}{3 X^{2}}\left(\sigma^{2}+\frac{3 X^{4} \sigma \partial_{\sigma} V}{\partial_{\eta}^{2} V}\right)^{1 / 2}, & f_{2}=\frac{X^{2} \partial_{\sigma} V \partial_{\eta}^{2} V}{3 \Lambda}, \quad f_{3}=\frac{X^{2} \partial_{\eta}^{2} V}{3 \sigma \partial_{\sigma} V}  \tag{5.3}\\
f_{6}=(18)^{2} \frac{3 X^{4}\left(\sigma^{2} \partial_{\sigma} V\right)\left(\partial_{\eta}^{2} V\right)}{\left(3 X^{4} \partial_{\sigma} V+\sigma \partial_{\eta}^{2} V\right)^{2}} \Lambda, & f_{7}=18\left(\partial_{\eta} V+\frac{3 X^{4} \sigma \partial_{\sigma} V \partial_{\sigma \eta}^{2} V}{3 X^{4} \partial_{\sigma} V+\sigma \partial_{\eta}^{2} V}\right), \\
f_{4}=\frac{2}{9}\left(\eta-\frac{\sigma \partial_{\sigma} V \partial_{\sigma \eta}^{2} V}{\Lambda}\right), & f_{5}=4\left(V-\frac{\sigma \partial_{\sigma} V\left(\partial_{\eta} V \partial_{\sigma \eta}^{2} V-3 X^{4} \partial_{\eta}^{2} V \partial_{\sigma} V\right)}{\Lambda}\right), \\
\Lambda=3 X^{4} \partial_{\eta}^{2} V \partial_{\sigma} V+\sigma\left[\left(\partial_{\eta \sigma}^{2} V\right)^{2}+\left(\partial_{\eta}^{2} V\right)^{2}\right], &
\end{array}
$$

and reduce to them for $X=1$. The sphere $\tilde{S}^{2}$ is fibered over the six-dimensional spacetime,

$$
\begin{equation*}
\operatorname{Vol}\left(\tilde{S}^{2}\right)=\epsilon^{i j k} y^{i} D y^{j} \wedge D y^{k}, \quad d s_{\tilde{S}^{2}}^{2}=D y^{i} D y^{i} . \tag{5.4}
\end{equation*}
$$

Here, $y^{i}$ are the embedding coordinate of the $S^{2}$ which can be chosen as

$$
\begin{equation*}
y^{1}=\sin \theta \sin \varphi, \quad y^{2}=\sin \theta \cos \varphi, \quad y^{3}=-\cos \theta, \tag{5.5}
\end{equation*}
$$

The symbol $D$ denotes the covariant derivative, as defined in eq. (2.10). Finally, $G_{5}$ is a differential form defined as
$G_{5}=-\frac{4}{9} \tilde{g}^{2} \sigma X^{4} \partial_{\eta}^{2} V *_{6} F_{3} \wedge d \eta \wedge d \sigma+\frac{2}{3 X^{2}} \tilde{g}\left(*_{6} F_{2}\right) \wedge *_{2} d\left(\sigma^{2} \partial_{\sigma} V\right)-\frac{2}{3 X^{2}}\left(*_{6} F^{i}\right) \wedge D\left(y^{i} \sigma^{2} \partial_{\sigma} V\right)$.
Let us make a quick summary. Any six dimensional configuration solution of the equation of motion (2.6)-(2.11) can be lifted to type IIB as in eqs. (5.1)-(5.6) and solves the equations of motion. Physically, the real scalar $X$ deforms the dilaton, axion and the warp factors as can be read in eqs. (5.3). The six dimensional fields $F_{2}, F_{3}, F^{i}$ enter the expressions of the fluxes using eqs. (5.1), (5.4) and (5.6) and fibre the two-sphere with the six dimensional space as expressed by eq. (5.4).

### 5.1 The explicit backgrounds

Let us now explicitly write the uplift of the solution presented in section 3 to type IIB supergravity. For each of these backgrounds, the ten dimensional supergravity equations of motion and Bianchi identities have been checked at the AdS-fixed points.

## 5.2 $\mathrm{AdS}_{4} \times \mathrm{H}_{2}$ in Type IIB

We write here the background in Type IIB, describing the flow from $\mathrm{AdS}_{6}$ to $\mathrm{AdS}_{4} \times H_{2}$, as found in section 3.1. We use the lift as described in eqs. (5.1)-(5.6).

$$
\begin{align*}
d s_{s t}^{2}=f_{1}[ & \frac{2 \tilde{g}^{2}}{9}\left(e^{2 f(r)}\left(-d t^{2}+d y_{1}^{2}+d y_{2}^{2}+e^{2 g(r)} d r^{2}\right)+e^{2 h(r)} \frac{\left(d z^{2}+d x^{2}\right)}{z^{2}}\right) \\
& \left.+f_{2} d s^{2}\left(\tilde{S}^{2}\right)+f_{3}\left(d \sigma^{2}+d \eta^{2}\right)\right] \\
B_{2}= & f_{4} \operatorname{Vol}\left(\tilde{S}^{2}\right)+\frac{2}{9} \eta \frac{\cos \theta}{z^{2}} d x \wedge d z, \quad C_{2}=f_{5} \operatorname{Vol}\left(\tilde{S}^{2}\right)+4 \partial_{\sigma}(\sigma V) \frac{\cos \theta}{z^{2}} d x \wedge d z \\
G_{5}= & \frac{2}{3 X^{2}}\left(\frac{\sqrt{2} \tilde{g}}{3}\right)^{4} e^{4 f(r)-2 h(r)+g(r)} d r \wedge d t \wedge d y_{1} \wedge d y_{2} \wedge d\left(\cos \theta \sigma^{2} \partial_{\sigma} V\right) \\
F_{5}= & 4\left(G_{5}+*_{10} G_{5}\right), \quad C_{0}=f_{7}, \quad e^{-2 \Phi}=f_{6} \tag{5.7}
\end{align*}
$$

We have used,

$$
\begin{equation*}
d s^{2}\left(\tilde{S}^{2}\right)=d \theta^{2}+\sin ^{2} \theta\left(d \phi-A^{3}\right)^{2}, \quad \operatorname{Vol}\left(\tilde{S}^{2}\right)=\sin \theta d \theta \wedge\left(d \phi-A^{3}\right), \quad A^{3}=\frac{d x}{z} \tag{5.8}
\end{equation*}
$$

We also used the definitions of $f_{1}, \ldots, f_{7}$ in eq. (5.3). These depend explicitly on the field $X(r)$. The expressions for $X(r), f(r), h(r)$ and $g(r)$ are read from section 3.1 and appendix B.1.

When considered at the fixed point of eq. (3.6), the configuration of eq. (5.7) solves all the equations of motion and describes an infinite family of $\mathrm{AdS}_{4}$ backgrounds in Type IIB preserving four Poincaré supercharges. This family is labelled by a function $V(\sigma, \eta)$ solving eq. (4.7). These solutions should be part of the more generic backgrounds in [50]. At the fixed point, the backgrounds in eq (5.7) are dual to an infinite family of three dimensional $\mathcal{N}=2$ SCFTs.

Let us now move into a new family of backgrounds with an $\mathrm{AdS}_{3} \times H_{3}$ factor in Type IIB.

## 5.3 $\mathrm{AdS}_{3} \times \mathrm{H}_{3}$ in Type IIB

This case is a bit more complicated than the previous one. Below, we write the background in Type IIB, describing the flow from $\mathrm{AdS}_{6}$ to $\mathrm{AdS}_{3} \times H_{3}$, as found in section 3.2. As above,
we use the lift as described in eqs. (5.1)-(5.6).

$$
\begin{align*}
d s_{s t}^{2}= & f_{1}\left[\frac{2 \tilde{g}^{2}}{9}\left(e^{2 f(r)}\left(-d t^{2}+d y^{2}+e^{2 g(r)} d r^{2}\right)+e^{2 h(r)} \frac{\left(d x_{1}^{2}+d x_{2}^{2}+d z^{2}\right)}{z^{2}}\right)\right. \\
& \left.+f_{2} d s^{2}\left(\tilde{S}^{2}\right)+f_{3}\left(d \sigma^{2}+d \eta^{2}\right)\right] \\
B_{2}= & f_{4} \operatorname{Vol}\left(\tilde{S}^{2}\right)-\frac{2}{9} \eta y^{i} F^{i}, C_{2}=f_{5} \operatorname{Vol}\left(\tilde{S}^{2}\right)-4 \partial_{\sigma}(\sigma V) y^{i} F^{i} \\
G_{5}= & -\frac{2}{3 X^{2}}\left(\frac{\sqrt{2} \tilde{g}}{3}\right)^{2} \frac{e^{3 f(r)-h(r)+g(r)}}{z} d r \wedge d t \wedge d y_{1} \wedge\left[d x_{2} \wedge d\left(\sin \theta \sin \phi \sigma^{2} \partial_{\sigma} V\right)\right. \\
& \left.+d x_{1} \wedge d\left(\cos \theta \sigma^{2} \partial_{\sigma} V\right)-d z \wedge\left(d\left(\cos \phi \sin \theta \sigma^{2} \partial_{\sigma} V\right)+\sigma^{2} \partial_{\sigma} V\left(\cos \theta A_{1}+\sin \theta \sin \phi A_{2}\right)\right)\right], \\
F_{5}= & 4\left(G_{5}+*_{10} G_{5}\right), C_{0}=f_{7}, e^{-2 \Phi}=f_{6} . \tag{5.9}
\end{align*}
$$

We have used the definitions in eqs. (3.10)-(5.5) and

$$
\begin{align*}
d s^{2}\left(\tilde{S}^{2}\right) & =\left(d \theta+\cos \phi A^{1}\right)^{2}+\sin ^{2} \theta\left(d \phi-\cot \theta \sin \phi A^{1}-A^{3}\right)^{2},  \tag{5.10}\\
\operatorname{Vol}\left(\tilde{S}^{2}\right) & =\sin \theta\left(d \theta+\cos \phi A^{1}\right) \wedge\left(d \phi-\cot \theta \sin \phi A^{1}-A^{3}\right) \\
A^{1} & =-\frac{d x_{1}}{z}, A^{2}=0, A^{3}=-\frac{d x_{2}}{z}, F^{1}=\frac{d z \wedge d x_{1}}{z^{2}}, F^{2}=\frac{d x_{2} \wedge d x_{1}}{z^{2}}, F^{3}=\frac{d z \wedge d x_{2}}{z^{2}},
\end{align*}
$$

The functions $X(r), f(r), g(r), h(r)$ solve the equations discussed in section 3.2 and appendix B.2. For the particular values in eq. (3.14) we checked that the Type IIB equations of motion are satisfied. In this case the background in eq. (5.9) describes an infinite family of Type IIB SUSY backgrounds with an $\mathrm{AdS}_{3}$ factor. There are dual to two dimensional $\mathcal{N}=(1,1)$ CFTs.

With the mechanism of the lift probably clear in the reader's mind, let us briefly discuss the new family of Type IIB backgrounds with an $\mathrm{AdS}_{2}$ factor.

## 5.4 $\mathrm{AdS}_{2} \times \mathrm{H}_{2} \times \mathrm{H}_{2}$ in Type IIB

The details needed to explicitly write the ten-dimensional solution interpolating between $\mathrm{AdS}_{6}$ and $\mathrm{AdS}_{2}$ are the definitions of the fields $A^{i}$ in eq. (3.18), the line and volume element for the two sphere as written below. With these elements, we carefully follow eqs. (5.1)(5.6) and write the new infinite family of Type IIB solutions with an $\mathrm{AdS}_{2}$ factor. This family of backgrounds reads,

$$
\begin{align*}
d s_{s t}^{2}=f_{1}[ & \frac{2 \tilde{g}^{2}}{9}\left(e^{2 f(r)}\left(-d t^{2}+e^{2 g(r)} d r^{2}\right)+e^{2 h_{1}(r)} \frac{\left(d z_{1}^{2}+d x_{1}^{2}\right)}{z_{1}^{2}}+e^{2 h_{2}(r)} \frac{\left(d z_{2}^{2}+d x_{2}^{2}\right)}{z_{2}^{2}}\right) \\
& \left.+f_{2} d s^{2}\left(\tilde{S}^{2}\right)+f_{3}\left(d \sigma^{2}+d \eta^{2}\right)\right] \\
B_{2}= & f_{4} \operatorname{Vol}\left(\tilde{S}^{2}\right)-\frac{2 \tilde{g}}{9} \sigma F_{2}+\frac{2}{9} \eta \cos \theta F^{3}, C_{2}=f_{5} \operatorname{Vol}\left(\tilde{S}^{2}\right)-4 \tilde{g}\left(\sigma \partial_{\eta} V\right) F_{2}+4 \partial_{\sigma}(\sigma V) \cos \theta F^{3} \\
G_{5}= & -\left(\frac{\sqrt{2}}{3} \tilde{g}\right)^{2}\left(\frac{1}{2 \tilde{g}^{2}} F^{3} \wedge F^{3} \wedge *_{2} d\left(\sigma^{2} \partial_{\sigma} V\right)+\frac{2 e^{2 f}}{3 X^{2}} d r \wedge d t \wedge F^{3} \wedge d\left(\cos \theta \sigma^{2} \partial_{\sigma} V\right)\right) \\
F_{5}= & 4\left(G_{5}+*_{10} G_{5}\right), \quad C_{0}=f_{7}, \quad e^{-2 \Phi}=f_{6} \tag{5.11}
\end{align*}
$$

We have used that,

$$
\begin{aligned}
A^{3} & =\frac{d x_{1}}{z_{1}}+\frac{d x_{2}}{z_{2}}, & F^{3}=-\frac{d z_{1} \wedge d x_{1}}{z_{1}^{2}}-\frac{d z_{2} \wedge d x_{2}}{z_{2}^{2}}, \\
F_{2} & =\frac{3}{2 \tilde{g}^{3}} X(r)^{2} e^{2\left(f(r)-h_{1}(r)-h_{2}(r)\right)} d t \wedge d r, & \\
d s^{2}\left(\tilde{S}^{2}\right) & =d \theta^{2}+\sin ^{2} \theta\left(d \phi-A^{3}\right)^{2}, & \operatorname{Vol}\left(\tilde{S}^{2}\right)=\sin \theta d \theta \wedge\left(d \phi-A^{3}\right) .
\end{aligned}
$$

In the case in which the fields $X(r), e^{2 f(r)}, e^{2 h_{1}(r)}$ and $e^{2 h_{2}(r)}$ take the values in eq. (3.25) we have checked that the configuration in eq. (5.11) solves the Type IIB equations of motion at the $\mathrm{AdS}_{2}$ fixed point. In eq. (5.11), we have an infinite family of $\mathrm{AdS}_{2}$ backgrounds in type IIB subject to the function $V(\sigma, \eta)$ solving the PDE (4.7).

## $5.5 \quad R^{1,3} \times S^{1}$ in Type IIB

The lift to Type IIB of the six dimensional background in section 3.4 is specially simple. In fact, the function $X(r)$ - the scalar in Romans supergravity is $X=1$. The fields $A_{2}, A_{1}$ and $A^{i}$ are all vanishing. As a consequence of this, the line and volume element of $\tilde{S}^{2}$ are those of the rounded sphere, the Ramond five form vanish and we have,

$$
\begin{align*}
d s_{s t}^{2} & =f_{1}\left[\frac{2 \tilde{g}^{2}}{9}\left(e^{2 \rho}\left(-d \tau^{2}+d \vec{x}_{3}^{2}\right)+\frac{d \rho^{2}}{g(\rho)}+e^{2 \rho} g(\rho) d \psi^{2}\right)+f_{2} d s^{2}\left(S^{2}\right)+f_{3}\left(d \sigma^{2}+d \eta^{2}\right)\right], \\
B_{2} & =f_{4} \operatorname{Vol}\left(S^{2}\right), \quad C_{2}=f_{5} \operatorname{Vol}\left(S^{2}\right), \quad C_{0}=f_{7}, \quad e^{-2 \Phi}=f_{6} . \tag{5.12}
\end{align*}
$$

The function $g(\rho)=1-e^{5\left(\rho_{*}-\rho\right)}$ with the constant $e^{2 \rho_{*}}=\frac{4}{125}$ to avoid conical singularities. The functions $f_{i}(\sigma, \eta)$ are identical to the $\tilde{f}_{i}(\sigma, \eta)$ in (4.8), which in turns are equivalent to eq. (5.3) after setting $X=1$. All the functions $V(\sigma, \eta)$ solving eq. (4.7) define a solution in Type IIB with a four dimensional Minkowski space-time and a circle (breaking all SUSYs) that asymptotes to $\mathrm{AdS}_{6} \times S^{2} \times \Sigma_{2}(\sigma, \eta)$. It is possible to construct another family of solutions using the non-SUSY $\operatorname{AdS}_{6}$ vacuum with $3 X^{4}=1$. The stability of this can be tested using the methods of [64].

## 6 Dual field theories

In this section we discuss aspects of the field theories dual to the type IIB backgrounds in sections $5.2-5.5$. We briefly remind the reader some aspects of 'twisted compactifications'. After that we discuss a meaningful observable, the holographic central charge. We define this quantity for the fixed-point solutions. Then, we extend this definition for a quantity along the full flows from $\mathrm{AdS}_{6}$.

### 6.1 Field theory duals to the Type IIB backgrounds

The backgrounds in sections 5.2-5.5 have holographic duals. The logic to find these duals goes as follows. In a certain regime of the radial-coordinate $r$ (or $\rho$ in section 5.5), the background asymptotes to $\mathrm{AdS}_{6}$. The field $X(r) \sim 1$ in that regime, but we still have the fields $F^{i}$ (and $F_{2}$ in section 5.4 ) switched on, breaking the $\operatorname{SO}(2,5)$ isometry. From the ten
dimensional perspective the $\mathrm{AdS}_{6}$ isometries are broken both with the fluxes and the metric, which has a non trivial $S^{2}$ fibration. This is characteristic of twisted compactifications. In this case, an infinite family of 5 d SCFTs (described by all the functions $V(\sigma, \eta)$ solving a Laplace equation) are compactified on curved manifolds: a hyperbolic plane $H_{2}$ in the case of section 5.2, a hyperbolic three space $H_{3}$ in section 5.3, and a product $H_{2} \times H_{2}$ in section 5.4. The twisted compactifications were introduced by Witten and nicely reviewed in [65]. In the context of D-branes the idea was developed by Berdshadsky, Sadov and Vafa in [66]. Holographically, these compactifications were described in [67] and then worked out in detail in many examples, see [58, 68-75] for a very small sample of such studies. For compactifications closely related to those discussed in sections $3.1-3.3$, see [76, 77].

In the case at hand, we think our backgrounds as holographically describing the twisted compactification of five dimensional SCFTs. A case with good analytical control is the one of balanced linear quivers [42]. To be concrete, we stick to this case in what follows, but there results will apply for more generic SCFTs. These five dimensional linear quivers, reach a conformal fixed point at high energies. The SCFT is deformed by operators describing the twisted compactification on a curved manifold ( $\mathrm{H}_{2}, \mathrm{H}_{3}$ and $\mathrm{H}_{2} \times \mathrm{H}_{2}$ in the cases mentioned above). The non-trivial prediction of our geometries is that at low energies compared with the finite-size of the compact space, the field theories flow to an interacting super conformal field theory in dimensions $(2+1),(1+1)$ and $(0+1)$ respectively. Also, that some fraction of the original eight Poincaré supercharges is preserved at low energies. Other non-trivial predictions of the type IIB backgrounds are explored in section 6.2.

The five dimensional SCFTs that we topologically-twist, describe the strongly coupled dynamics (at high energies) of a linear quiver field theory of the form


The numbers $N_{1}, N_{2}, \ldots, N_{P}$ and $F_{1}, \ldots, F_{P}{ }^{3}$ determine uniquely the function $V(\sigma, \eta)$. This is clearly explained in [42], we refer the reader to that work.

These UV conformal points are deformed by relevant operators. The dimension of these operators can be read from the near- $\mathrm{AdS}_{6}$ expansion of the gauged supergravity metric. These relevant operators (analogously the presence of the six dimensional gauge fields) topologically twist the 5d CFT and trigger a RG flow, that ends in a $\mathrm{CFT}_{3}, \mathrm{CFT}_{2}$, $\mathrm{CFT}_{1}$ for the backgrounds in sections $5.2,5.3,5.4$ respectively. An interesting quantity measuring the number of degrees of freedom (or the Free Energy) of these strongly coupled lower dimensional CFTs is defined in section 6.2.1. As we discuss below, the result is in terms of transcendental functions of the parameters defining the quiver, hence revealing the non-perturbative character of the CFT. An interesting result is that, in terms of the

[^2]quiver parameters, the holographic central charge of the IR CFT is proportional to the UV one. This supports the picture advocated by Bobev-Crichigno [58].

In contrast, the solution of section 5.5 can be obtained by Wick rotating a Schwarzchild black hole in $\mathrm{AdS}_{6}$. This background is also called the AdS-soliton. The solution in section 5.5, describes a family of five dimensional SCFTs, that at the UV fixed point are compactified on a circle. After imposing periodic boundary conditions for the bosons and anti-periodic ones for the fermions (along a spatial circle), supersymmetry is broken, turning the scalars and fermions massive. In the perturbative spectrum, the gauge fields are massless. Our family of backgrounds describe the strong-dynamics of these systems.

Similar dynamics was exploited in [78]-[81] with phenomenological purposes. It would be of interest to determine if the presence of bifundamental matter (even when massive) introduces any novel behaviour on observables.

### 6.2 The holographic central charge

In this section we calculate a quantity called holographic central charge. We start in section 6.2 .1 by defining the quantity at conformal points and explicitly computing it in our $\operatorname{AdS}_{4}, \mathrm{AdS}_{3}$ and $\mathrm{AdS}_{2}$ fixed point type IIB geometries. This gives us information about the number of degrees of freedom in the strongly coupled $\mathrm{CFT}_{3}, \mathrm{CFT}_{2}$ and $\mathrm{CFT}_{1}$ respectively.

After that, in section 6.2 .2 we present a quantity originally defined in [82]. This quantity is inspired on the holographic central charge of [83], but is applicable to geometries describing flows, like our $\mathrm{AdS}_{6} \rightarrow \mathrm{AdS}_{4,3,2}$. We call this quantity the 'flow central charge'. The main characteristic is that it is constant at both ends of the flow. As we explicitly show for our solutions, this quantity is monotonous along the flow and its value at the IR-fixed point is bigger than that at the UV-fixed point.

Finally, in section 6.2 .3 we compute the flow-central charge for the solution in section 5.5 and interpret its result physically.

### 6.2.1 The holographic central charge: definition and calculation at fixed points

First, we briefly remind the reader the formalism we use, developed in [82, 83]. It can be briefly summarised as follows: consider a $(d+1)$ dimensional QFT dual to a background, with metric and dilaton of the form,

$$
\begin{equation*}
d s^{2}=a(r, \vec{\theta})\left[-d t^{2}+d \vec{x}_{d}^{2}+b(r) d r^{2}\right]+g_{i j}(r, \vec{\theta}) d \theta^{i} d \theta^{j}, \quad e^{\Phi(r, \vec{\theta}} . \tag{6.1}
\end{equation*}
$$

We calculate a weighted version of the internal volume,

$$
\begin{equation*}
V_{\mathrm{int}}(r)=\int d \vec{\theta} \sqrt{e^{-4 \Phi(r, \vec{\theta})} \operatorname{det}\left[g_{i j}\right] a^{d}(r, \vec{\theta})} . \tag{6.2}
\end{equation*}
$$

We define then a quantity $H(r)=V_{\text {int }}^{2}$ and compute the holographic central charge according to

$$
\begin{equation*}
c_{\mathrm{hol}}=\frac{d^{d}}{G_{N}} b(r)^{\frac{d}{2}} \frac{H^{\frac{2 d+1}{2}}}{\left(H^{\prime}\right)^{d}} . \tag{6.3}
\end{equation*}
$$

Let us see this at work. First consider the fixed point solutions with $\mathrm{AdS}_{4}$, then $\mathrm{AdS}_{3}$ and finally $\mathrm{AdS}_{2}$. In these three fixed point solutions the full space time metric and dilaton are,

$$
\begin{equation*}
d s_{10, s t}^{2}=f_{1}(\sigma, \eta)\left[\left(\frac{2 \tilde{g}^{2}}{9}\right) d s_{6}^{2}+f_{2}(\sigma, \eta) d \tilde{\Omega}_{2}+f_{3}(\sigma, \eta)\left(d \sigma^{2}+d \eta^{2}\right)\right], \quad e^{-2 \phi}=f_{6}(\sigma, \eta) \tag{6.4}
\end{equation*}
$$

For each of the fixed points we have

$$
\begin{array}{ll}
d s_{6, \mathrm{AdS}_{4}}^{2}=e^{2 f(r)}\left(-d t^{2}+d y_{1}^{2}+d y_{2}^{2}+e^{2 g(r)} d r^{2}\right)+\frac{e^{2 h(r)}}{z^{2}}\left(d z^{2}+d x^{2}\right), & d=2, \\
d s_{6, \mathrm{AdS}_{3}}^{2}=e^{2 f(r)}\left(-d t^{2}+d y_{1}^{2}+e^{2 g(r)} d r^{2}\right)+\frac{e^{2 h(r)}}{z^{2}}\left(d z^{2}+d x_{1}^{2}+d x_{2}^{2}\right), & d=1, \\
d s_{6, \mathrm{AdS}_{2}}^{2}=e^{2 f(r)}\left(-d t^{2}+e^{2 g(r)} d r^{2}\right)+\frac{e^{2 h_{1}(r)}}{z_{1}^{2}}\left(d z_{1}^{2}+d x^{2}\right)+\frac{e^{2 h_{2}(r)}}{z_{2}^{2}}\left(d z_{2}^{2}+d x^{2}\right), & d=0 \tag{6.7}
\end{array}
$$

The functions

$$
\begin{equation*}
a(r, \vec{\theta})=\frac{2 \tilde{g}^{2}}{9} f_{1}(\sigma, \eta) e^{2 f(r)}, \quad b(r)=e^{2 g(r)} \tag{6.8}
\end{equation*}
$$

The 'internal spaces' are

$$
\begin{align*}
& d s_{\mathrm{int}, \mathrm{AdS}_{4}}^{2}=f_{1}\left[\frac{2 \tilde{g}^{2}}{9} \frac{e^{2 h(r)}}{z^{2}}\left(d z^{2}+d x^{2}\right)+f_{2} d \tilde{\Omega}_{2}+f_{3}\left(d \sigma^{2}+d \eta^{2}\right)\right],  \tag{6.9}\\
& d s_{\mathrm{int}, \mathrm{AdS}_{3}}^{2}=f_{1}\left[\frac{2 \tilde{g}^{2}}{9} \frac{e^{2 h(r)}}{z^{2}}\left(d z^{2}+d x_{1}^{2}+d x_{2}^{2}\right)+f_{2} d \tilde{\Omega}_{2}+f_{3}\left(d \sigma^{2}+d \eta^{2}\right)\right],  \tag{6.10}\\
& d s_{\mathrm{int}, \mathrm{AdS}_{2}}^{2}=f_{1}\left[\frac{2 \tilde{g}^{2}}{9} \frac{e^{2 h_{1}(r)}}{z_{1}^{2}}\left(d z_{1}^{2}+d x_{1}^{2}\right)+\frac{e^{2 h_{2}(r)}}{z_{2}^{2}}\left(d z_{2}^{2}+d x_{2}^{2}\right)+f_{2} d \tilde{\Omega}_{2}+f_{3}\left(d \sigma^{2}+d \eta^{2}\right)\right] . \tag{6.11}
\end{align*}
$$

Calculating the combination in eq. (6.2) we have,

$$
\begin{align*}
& V_{\mathrm{int}, \mathrm{AdS}_{4}}=\left(\operatorname{Vol}\left(S^{2}\right) \operatorname{Vol}\left(H_{2}\right)\left(\frac{2 \tilde{g}^{2}}{9}\right)^{2} \int d \sigma d \eta f_{1}^{4} f_{2} f_{3} f_{6}\right) \times e^{2 f(r)+2 h(r)},  \tag{6.12}\\
& V_{\mathrm{int}, \mathrm{AdS}_{3}}=\left(\operatorname{Vol}\left(S^{2}\right) \operatorname{Vol}\left(H_{3}\right)\left(\frac{2 \tilde{g}^{2}}{9}\right)^{2} \int d \sigma d \eta f_{1}^{4} f_{2} f_{3} f_{6}\right) \times e^{f(r)+3 h(r)},  \tag{6.13}\\
& V_{\mathrm{int}, \mathrm{AdS}_{2}}=\left(\operatorname{Vol}\left(S^{2}\right) \operatorname{Vol}\left(H_{2,1}\right) \operatorname{Vol}\left(H_{2,2}\right)\left(\frac{2 \tilde{g}^{2}}{9}\right)^{2} \int d \sigma d \eta f_{1}^{4} f_{2} f_{3} f_{6}\right) \times e^{2 h_{1}(r)+2 h_{2}(r)} \tag{6.14}
\end{align*}
$$

Using eq. (5.3), one finds

$$
\int d \sigma d \eta f_{1}^{4} f_{2} f_{3} f_{6}=\frac{2^{6}}{3} \int d \sigma d \eta \sigma^{3}\left(\partial_{\sigma} V\right)\left(\partial_{\eta}^{2} V\right)
$$

Now, if we define

$$
\begin{align*}
& \mathcal{N}_{\mathrm{AdS}_{4}}=\frac{2^{6}}{3} \operatorname{Vol}\left(S^{2}\right) \operatorname{Vol}\left(H_{2}\right)\left(\frac{2 \tilde{g}^{2}}{9}\right)^{2} \int d \sigma d \eta \sigma^{3}\left(\partial_{\sigma} V\right)\left(\partial_{\eta}^{2} V\right)  \tag{6.15}\\
& \mathcal{N}_{\mathrm{AdS}_{3}}=\frac{2^{6}}{3} \operatorname{Vol}\left(S^{2}\right) \operatorname{Vol}\left(H_{3}\right)\left(\frac{2 \tilde{g}^{2}}{9}\right)^{2} \int d \sigma d \eta \sigma^{3}\left(\partial_{\sigma} V\right)\left(\partial_{\eta}^{2} V\right)  \tag{6.16}\\
& \mathcal{N}_{\mathrm{AdS}_{2}}=\frac{2^{6}}{3} \operatorname{Vol}\left(S^{2}\right) \operatorname{Vol}\left(H_{2,1}\right) \operatorname{Vol}\left(H_{2,2}\right)\left(\frac{2 \tilde{g}^{2}}{9}\right)^{2} \int d \sigma d \eta \sigma^{3}\left(\partial_{\sigma} V\right)\left(\partial_{\eta}^{2} V\right) \tag{6.17}
\end{align*}
$$

we then have:

$$
\begin{aligned}
V_{\mathrm{int}, \mathrm{AdS}_{4}} & =\mathcal{N}_{\mathrm{AdS}_{4}} e^{2 f+2 h}, & V_{i n t, \mathrm{AdS}_{3}}=\mathcal{N}_{\mathrm{AdS}_{3}} e^{f+3 h}, & V_{\mathrm{int}^{2}, \mathrm{AdS}_{2}}=\mathcal{N}_{\mathrm{AdS}_{2}} e^{2 h_{1}+2 h_{2}} \\
H_{\mathrm{AdS}_{4}} & =\mathcal{N}_{\mathrm{AdS}_{4}}^{2} e^{4 f+4 h}, & H_{\mathrm{AdS}_{3}}=\mathcal{N}_{\mathrm{AdS}_{3}}^{2} e^{2 f+6 h}, & H_{\mathrm{AdS}_{2}}=\mathcal{N}_{\mathrm{AdS}_{2}}^{2} e^{4 h_{1}+4 h_{2}}
\end{aligned}
$$

Using eq. (6.3) we find the holographic central charge for each fixed point to be,

$$
\begin{align*}
& c_{\mathrm{hol}, \mathrm{AdS}_{4}}=\frac{\mathcal{N}_{\mathrm{AdS}_{4}}}{4 G_{N}}\left(\frac{e^{f+g+h}}{f^{\prime}+h^{\prime}}\right)^{2}=\frac{\mathcal{N}_{\mathrm{AdS}_{4}}}{4 G_{N}}\left(\frac{e^{f(r)+g(r)}}{f^{\prime}(r)}\right)^{2} e^{2 h_{0}}  \tag{6.18}\\
& c_{{\mathrm{hol}, \mathrm{AdS}_{3}}}=\frac{\mathcal{N}_{\mathrm{AdS}_{3}}}{G_{N}}\left(\frac{e^{f+g+3 h}}{f^{\prime}+3 h^{\prime}}\right)=\frac{\mathcal{N}_{\mathrm{AdS}_{3}}}{G_{N}}\left(\frac{e^{f(r)+g(r)}}{f^{\prime}(r)}\right) e^{3 h_{0}}  \tag{6.19}\\
& c_{\mathrm{hol}, \mathrm{AdS}_{2}}=\frac{\mathcal{N}_{\mathrm{AdS}_{2}}}{G_{N}}\left(e^{2 h_{1,0}+2 h_{2,0}}\right) \tag{6.20}
\end{align*}
$$

We have denoted by $h_{0}$ the value of the warp function $h(r)$ at the fixed point, see eqs. (3.6), (3.14), (3.25) respectively. We used that, at the fixed point $h^{\prime}(r)=0$, $f(r)=-\log \frac{r}{r_{0}}, g(r)=0$ (or we choose $g(r)=-f(r)=-\frac{r}{r_{0}}$ ). The value of $r_{0}$ can be computed using eqs. (3.6), (3.14). For the particular case of $\mathrm{AdS}_{2}$ with $d=0$ as indicated in eq. (6.7) it is enough to compute with $V_{\text {int }}$ in eq. (6.14). This subtlety was discussed in [2].

Note that the factors $\mathcal{N}_{\mathrm{AdS}_{4,3,2}}$ defined in eqs. (6.15)-(6.17) involve, aside form the volumes of the compact internal space, an integral,

$$
\begin{equation*}
\mathcal{N}_{\mathrm{AdS}_{6}}=\frac{2^{6}}{3} \operatorname{Vol}\left(S^{2}\right) \int d \sigma d \eta \sigma^{3}\left(\partial_{\sigma} V\right)\left(\partial_{\eta}^{2} V\right) \tag{6.21}
\end{equation*}
$$

This integral was found in [42] when computing the holographic central charge/Free Energy of a general 5d SCFTs (our far UV SCFTs), see equation (3.7) in [42]. In the case of balanced linear quivers, this five dimensional holographic central charge can be computed explicitly and it has a transcendental dependence on the parameters defining the 5 d quiver. See for example equation (3.21) in [42].

A way of physically understanding the expressions in eqs. (6.18)-(6.20) is to think that the degrees of freedom of the five dimensional UV conformal theory proportional to the quantity $\mathcal{N}_{\mathrm{AdS}_{6}}$ in eq. (6.21) gets 'weighted' by the volume of the compactification manifold, represented by $e^{h_{0}}, \operatorname{Vol}\left(H_{i}\right)$, etc. This confirms the picture advocated in [58].

To close this section and motivate the next one, we observe by inspecting the generic expressions in eqs. (6.18)-(6.19), that these do not present the fingerprint of a UV fixed
point. Namely, when $f(r) \sim h(r) \sim-\log r, g(r)=0$ we do not find a constant value for $c_{\text {hol }}$. In other words, the $\mathrm{AdS}_{6}$-UV fixed point of the flow is not captured by the quantity defined in eq. (6.3). We lost this fixed point when we informed the theory that the space-dimension of the dual CFT was either $d=2,1,0$ in eqs. (6.5)-(6.7). The quantity defined in the next section remedies this deficiency and shows the existence of both the IR and UV fixed points (except for the $\mathrm{AdS}_{2}$ as we discuss). We call this quantity the 'flow central charge'.

### 6.2.2 The flow-central charge: definition and calculations

The purpose of this section is to test a quantity defined in [82]. This quantity here named "flow central charge" should be monotonous (at least on the BPS flows), capture the fact that there is a CFT both at the beginning and at the end of the flow. The case of a mass gapped theory is discussed separately.

This quantity should not be thought as a measure of the number of degrees of freedom. Even when presenting a similar structure to the holographic central charge of section 6.2.1, it does not give the same result in the IR fixed points as those in the previous section.

An interesting feature of the flow central charge is that it separates the dynamics of the flow from the degrees of freedom of the UV SCFT. This is not surprising, given how these solutions have been constructed. There could be other quantities that present similar characteristics.

For the situation in which the QFT is actually a flow across dimensions (or the field theory lives in an anisotropic space time), we appeal to the slightly more elaborated formalism developed in [82]. Adapting the formalism of [82] to our flows from $\mathrm{AdS}_{6}$ backgrounds, leads us to consider metrics and dilaton of the form,

$$
\begin{align*}
d s^{2}= & -\alpha_{0} d t^{2}+\alpha_{1} d y_{1}^{2}+\alpha_{2} d y_{2}^{2}+\ldots+\alpha_{d} d y_{d}^{2}+\Pi_{i=1}^{d}\left(\alpha_{1} \ldots \alpha_{d}\right)^{\frac{1}{d}} b(r) d r^{2} \\
& +g_{i j}\left(d \theta^{i}-A_{1}^{i}\right)\left(d \theta^{j}-A_{1}^{j}\right), \quad e^{\Phi} . \tag{6.22}
\end{align*}
$$

For the cases studied here, we set $d=4$ as we perform flows from $\mathrm{AdS}_{6}$. We are interested in defining a quantity that is monotonous along the flow. We define,

$$
\begin{equation*}
d s_{\mathrm{int}}^{2}=\alpha_{1} d y_{1}^{2}+\alpha_{2} d y_{2}^{2}+\ldots+\alpha_{d} d y_{d}^{2}+g_{i j}\left(d \theta^{i}-A_{1}^{i}\right)\left(d \theta^{j}-A_{1}^{j}\right), \quad e^{\Phi} . \tag{6.23}
\end{equation*}
$$

We form the combination,

$$
\begin{equation*}
V_{\text {int }}=\int_{X} \sqrt{\operatorname{det}\left[g_{m n}\right] e^{-4 \Phi}}, \quad H=V_{\text {int }}^{2} . \tag{6.24}
\end{equation*}
$$

The integral is over $X$ the manifold consisting of the internal space $g_{i j}$ and the dimensions 'erased' by the RG-flow. Then we define the holographic central charge along the flow as in eq. (6.3). Namely, using that $d=4$ in all of our configurations,

$$
\begin{equation*}
c_{\text {flow }}=\frac{4^{4}}{G_{N}} b(r)^{2}\left(\frac{H}{H^{\prime}}\right)^{4} H^{1 / 2} . \tag{6.25}
\end{equation*}
$$

For the flow metric interpolating between $\mathrm{AdS}_{6}$ and $\mathrm{AdS}_{4} \times H_{2}$ in type IIB - see eqs. (5.7) and (6.5), we have

$$
\begin{array}{rlrl}
\alpha_{1} & =\alpha_{2}=\frac{2 \tilde{g}^{2}}{9} f_{1}(\sigma, \eta) e^{2 f(r)}, & \alpha_{3}=\alpha_{4}=\frac{2 \tilde{g}^{2}}{9} f_{1}(\sigma, \eta) e^{2 h(r)},  \tag{6.26}\\
b(r) & =e^{f(r)+2 g(r)-h(r)}, & V_{\mathrm{int}}=\mathcal{N}_{\mathrm{AdS}_{4}} e^{2 f(r)+2 h(r)}, & c_{\text {flow }}=\frac{\mathcal{N}_{\mathrm{AdS}_{4}}}{G_{N}}\left(\frac{e^{f+g}}{f^{\prime}+h^{\prime}}\right)^{4}
\end{array}
$$

We have used the definition for $\mathcal{N}_{\mathrm{AdS}_{4}}$ in eq. (6.15). For the IR fixed point, when $f=$ $-\log \frac{r}{r_{0}}, h=h_{0}, g=0$ and the UV fixed point with $f \sim h \sim-\log \frac{r}{r_{0}}$ and $g=0$ we find

$$
\begin{equation*}
c_{\mathrm{IR}, \text { flow }}=\frac{\mathcal{N}_{\mathrm{AdS}_{4}}}{G_{N}} r_{0}^{4}, \quad c_{\mathrm{UV}, \text { flow }}=\frac{c_{\mathrm{IR}, \text { flow }}}{16} \tag{6.27}
\end{equation*}
$$

As before, the number $r_{0}$ can be calculated by imposing $f(r)=-\log \frac{r}{r_{0}}$ and $g(r)=0$ in eq. (3.6). Notice that the value in the IR of the flow is proportional to that in eq. (6.18) and that the dependence on the 'internal space' expressed by $\mathcal{N}_{\mathrm{AdS}_{4}}$ is the same. This quantity does not indicate the number of degrees of freedom (as the IR of the flow contains more than the UV fixed point). But the quantity detects both fixed points.

The case of the flow $\mathrm{AdS}_{6} \rightarrow \mathrm{AdS}_{3} \times H_{3}$ in type IIB represented by eqs. (5.9), (6.6) works very much along the same lines, giving

$$
\begin{align*}
\alpha_{1} & =\frac{2 \tilde{g}^{2}}{9} f_{1}(\sigma, \eta) e^{2 f(r)}, & \alpha_{2} & =\alpha_{3}=\alpha_{4}=\frac{2 \tilde{g}^{2}}{9} f_{1}(\sigma, \eta) e^{2 h(r)}, \\
b(r) & =e^{\frac{3 f(r)}{2}+2 g(r)-\frac{3 h(r)}{2}}, & V_{\mathrm{int}} & =\mathcal{N}_{\mathrm{AdS}_{3}} e^{f(r)+3 h(r)}, \\
c_{\text {flow }} & =\frac{16 \mathcal{N}_{\mathrm{AdS} 3}}{G_{N}}\left(\frac{e^{f+g}}{f^{\prime}+3 h^{\prime}}\right)^{4}=\frac{16 \mathcal{N}_{\mathrm{AdS} 3}}{G_{N}} r_{0}^{4}, & c_{\mathrm{IR}, \text { flow }} & =256 c_{\mathrm{UV}, \text { flow }} \tag{6.28}
\end{align*}
$$

As above, the number $r_{0}$ can be calculated by imposing $f(r)=-\log \frac{r}{r_{0}}$ and $g(r)=0$ in eq. (3.14).

More interesting is the case of the flow in eqs. (5.11), (6.7). We obtain,

$$
\begin{array}{ll}
b(r) & =e^{2 f(r)+2 g(r)-h_{1}(r)-h_{2}(r)}, \\
c_{\text {flow }} & =\frac{\mathcal{N}_{\mathrm{AdS} 2}}{G_{N}}\left(\frac{e^{f+g}}{h_{1}^{\prime}+h_{2}^{\prime}}\right)^{4} \tag{6.29}
\end{array}
$$

This quantity detects the $\mathrm{UV}-\mathrm{AdS}_{6}$ fixed point, but becomes divergent in the IR fixed point, when $h_{1}=h_{2}=h_{0}$ becomes constant. Hence we see that this flow-central charge should be handled with care. For the BPS solutions shown in figures $1,2,3$, we plot the flow central charge (setting $G_{N}=1$ ). The result is displayed in figure 4.

In the next section, we study this same quantity for the flow of section 5.5 and discuss the implications in Physics.


Figure 4. Plot of the flow central charge for the BPS solutions discussed in figures 1,2 and 3 . We have chosen $G_{N}=1$. The monotonicity of this quantity is clearly displayed.

### 6.2.3 The flow-central charge for the $\operatorname{AdS}_{6} \rightarrow R^{1,3} \times S^{1}$ background

We now consider the family of backgrounds discussed in section 5.5. The metric and dilaton needed for this calculation are,

$$
\begin{aligned}
d s_{10, s t}^{2} & =f_{1}\left[e^{2 \rho}\left(-d \tau^{2}+d \vec{x}^{2}\right)+\frac{d \rho^{2}}{g(\rho)}+e^{2 \rho} g(\rho) d \psi^{2}+f_{2} d \Omega_{2}(\theta, \varphi)+f_{3}\left(d \sigma^{2}+d \eta^{2}\right)\right] \\
e^{-2 \Phi} & =972 e^{2 \Phi_{0}} \frac{\sigma^{2} \partial_{\sigma} V \partial_{\eta}^{2} V}{\left(3 \partial_{\sigma} V+\sigma \partial_{\eta}^{2} V\right)^{2}} \Lambda .
\end{aligned}
$$

The function $g(\rho)=1-e^{5\left(\rho_{*}-\rho\right)}$. The functions $f_{i}(\sigma, \eta)$ and $\Lambda(\sigma, \eta)$ can be read from eq. (4.8), or from eq. (5.3), after setting $X=1$.

We can view this background as describing a compactification on a circle or as a field theory that is five dimensional, but that has an anisotropy. We use the formula developed in [82], summarised in eqs. (6.22)-(6.25). To apply this requires the following assignation,

$$
\begin{align*}
d & =4, \quad b(\rho)=e^{-2 \rho} g(\rho)^{-\frac{5}{4}}  \tag{6.30}\\
d s_{\mathrm{int}}^{2} & =f_{1}\left[e^{2 \rho}\left(d x_{1}^{2}+d x_{2}^{2}+d x_{3}^{2}+g(\rho) d \psi^{2}\right)+f_{2} d \Omega_{2}+f_{3}\left(d \sigma^{2}+d \eta^{2}\right)\right] \\
V_{\mathrm{int}} & =\left(4 \pi \int d \sigma d \eta e^{-2 \Phi} f_{1}^{4} f_{2} f_{3}\right) e^{4 \rho} \sqrt{g(\rho)}=\mathcal{N} e^{4 \rho} \sqrt{g(\rho)} \\
H & =\mathcal{N}^{2} e^{8 \rho} g(\rho), \quad H^{\prime}=\mathcal{N}^{2} e^{8 \rho}\left(8 g(\rho)+g^{\prime}(\rho)\right) \\
c_{\mathrm{hol}} & =\frac{\mathcal{N}}{16 G_{N}} \frac{g^{2}}{\left(g+\frac{g^{\prime}}{8}\right)^{4}}=\frac{\mathcal{N}}{16 G_{N}} \frac{\left(1-e^{-5\left(\rho-\rho_{*}\right)}\right)^{2}}{\left(1-\frac{3}{8} e^{-5\left(\rho-\rho_{*}\right)}\right)^{4}}
\end{align*}
$$

This expression displays the expected behaviours. In the far IR, when $\rho \rightarrow \rho_{*}$, the value is zero indicating a gapped spectrum. At high energies, when $\rho \rightarrow \infty$, the result is that for a 5 d SCFT. In fact, compare with eqs. (3.6)-(3.7) of [42]. There is an infinite number of


Figure 5. Plot of the flow central charge for the non-SUSY solution discussed in section 5.5. We have chosen $G_{N}=1$. The monotonicity of this quantity is clearly displayed.
linear quivers in 5 d that upon compatification on a circle display a dynamics described by the background in section 5.5. These field theories differ in the choice of ranks of colour and flavour groups for the linear quiver that in the far UV reaches a conformal point.

Notice also the 'separation' between the flow - represented by the $\rho$-dependent part of the central charge, and the CFT structure -represented by $\frac{\mathcal{N}}{16 G_{N}}$ that this formula indicates. Similar 'decoupling phenomenon' in different flows were observed in [84]. Following the formalism in [85], we can calculate the Entanglement Entropy in this case and find a behaviour qualitatively similar to the one for the flow central charge. The flow central charge for this QFT is plotted in figure 5 .

## 7 Conclusions

Let us start with a brief summary of the idea and achievements of this paper and close with future lines of inquire.

After discussing four configurations in Romans' six dimensional $F_{4}$ gauged supergravity, we lift them to Type IIB generating new families of string backgrounds with $\mathrm{AdS}_{4}$, $\mathrm{AdS}_{3}$ and $\mathrm{AdS}_{2}$ that preserve SUSY. We also discussed a family of solutions describing a circle compactification $\mathrm{AdS}_{6} \rightarrow R^{1,3} \times S^{1} \times R_{\rho}$ that break SUSY. All these backgrounds are parameterized by a function $V(\sigma, \eta)$, solving a Laplace equation. The compactifications to lower dimensional $\mathrm{AdS}_{4,3,2}$ are dual to SCFTs in dimensions $(2+1),(1+1)$ and $(0+1)$ dimensions. We identify the dual field theory as the twisted compactification of a 5 d 'mother' SCFT on hyperbolic spaces of dimensions $2,3,4$ respectively. The 'mother' five dimensional SCFT preserves eight Poincaré SUSYs, it gets deformed by relevant operators that implement the twisted compactification. The kinematic data of the mother 5d SCFT is encoded in the function $V(\sigma, \eta)$ and the associated quiver and matter content is inherited by the lower dimensional IR SCFT.

We study the number of degrees of freedom/Free Energy of the lower dimensional fixed points and emphasise the non-perturbative nature of the result. We also define a monotonic quantity, characterising the flow. The non-SUSY compactification to a four dimensional quiver QFT with massive matter is also treated.

In the near future, it would be interesting to study different twisted compactifications and extend the picture displayed here to other Type II or M-theory backgrounds. Some steps in this direction are reported in [84]. The study of the Holographic Entanglement Entropy in these flow-geometries following the formalism of [85] might also give interesting information.

The development of a map between SCFTs and supergravity backgrounds with less than half-BPS SUSY preserved is a worthwhile project. In this vein, it would be interesting (but not easy) to cast our infinite family of background with an $\mathrm{AdS}_{4}$ factor in section 5.2 in the language of [50]. There should be a way of relating our family of solutions with that of [53]. The formalism developed in [86] should apply for our backgrounds in section 5.2. There may be similar field theory elaborations for our backgrounds in sections 5.3 and 5.4. Obviously, a better understanding of the field theory dual to the background in section 5.5 is very desirable, mostly to be used as a model with possible phenomenological interest.

We hope to report some progress in some of these lines above mentioned.

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## A Conventions

Our conventions agree with the one in [59] except for the definition of the Hodge dual of a $k$-form, which in our convention is given by

$$
\begin{equation*}
*\left(d x^{\mu_{1}} \wedge \cdots \wedge d x^{\mu_{k}}\right)=\frac{\sqrt{|g|}}{(D-k)} \epsilon^{\mu_{1} \ldots \mu_{k}}{ }_{\nu_{1} \ldots \nu_{D-k}} d x^{\nu_{1}} \wedge \cdots \wedge d x^{\nu_{D-k}} \tag{A.1}
\end{equation*}
$$

where $D$ is the space-time dimension and

$$
\begin{equation*}
\epsilon_{12 \ldots D}=1 . \tag{A.2}
\end{equation*}
$$

Notice that this definition differ respect to the one in [59] for the sign of the Hodge-dual of odd-forms; for example, we have

$$
\begin{equation*}
*_{2} D y^{i}=-\epsilon_{i j k} y^{j} D y^{k} . \tag{A.3}
\end{equation*}
$$

This lead also to different signs in the kinetic terms in (2.5) and in the equations of motion (2.6)-(2.9).

## B Numerics

In this section we give some details on the numerical analysis which leads to the flows from $\mathrm{AdS}_{6}$ to lower dimensional AdS solutions. For the $\mathrm{AdS}_{d}$ fixed points the idea will be the same: solve a linearized version of the BPS system close to the IR fixed point and use that analysis to give the initialization values for the numerical analysis. In the following, we will set $\tilde{g}=2 / \sqrt{2}$ and we fix the parameterization invariance by setting $g=-f+\log X$. We also define $h=-\frac{1}{2} \log Y$.

## B. $1 \quad$ AdS $_{4}$ flow

After the redefinition in the previous above the BPS system (3.3)-(3.5) reads

$$
\begin{align*}
X^{\prime}+\frac{3}{4} X\left(X^{-2}-X^{2}-\frac{Y}{9}\right) & =0,  \tag{B.1}\\
3 X^{2}+X^{-2}-\frac{2 Y^{\prime}}{Y}-Y & =0,  \tag{B.2}\\
12 f^{\prime}+9 X^{2}+3 X^{-2}+Y & =0 . \tag{B.3}
\end{align*}
$$

Using (B.1) as definition of $Y$, we can get a definition of $f$ in terms of $X$ and a second order ODE for $X$ :

$$
\begin{equation*}
f^{\prime}=-\frac{1}{X^{2}}-\frac{X^{\prime}}{X}, \quad 2 X^{\prime \prime}+\frac{10 X^{\prime 2}}{X}+\left(\frac{14}{X^{2}}-24 X^{2}\right) X^{\prime}+9 X^{5}-15 X+\frac{6}{X^{3}}=0 \tag{B.4}
\end{equation*}
$$

We can now solve this equation linearizing it close to the fixed point:

$$
\begin{equation*}
X=\sqrt[4]{\frac{2}{3}}+\epsilon x(r) \tag{B.5}
\end{equation*}
$$

At the first order in $\epsilon, x$ has to solve

$$
\begin{equation*}
2 x^{\prime \prime}-\sqrt{6} x^{\prime}-12 x=0 \tag{B.6}
\end{equation*}
$$

which admits the solution

$$
\begin{equation*}
x=e^{-\frac{1}{2} \sqrt{\frac{3}{2}}(\sqrt{17}-1) r}\left(c_{2} e^{\sqrt{\frac{51}{2}} r}+c_{1}\right) \tag{B.7}
\end{equation*}
$$

Using this definition of $X$ we can derive the ones of $f$ and $Y$ at the first order in $\epsilon$ :

$$
\begin{align*}
Y & =3 \sqrt{\frac{3}{2}}+\epsilon 3 \sqrt[4]{54} e^{-\frac{1}{2} \sqrt{\frac{3}{2}}(\sqrt{17}-1) r}\left((\sqrt{17}-4) c_{2} e^{\sqrt{\frac{51}{2}} r}-(\sqrt{17}+4) c_{1}\right)  \tag{B.8}\\
f & =-\sqrt{\frac{3}{2}} r+\epsilon \frac{1}{4} \sqrt[4]{\frac{3}{2}} e^{-\frac{1}{2} \sqrt{\frac{3}{2}}(\sqrt{17}-1) r}\left((\sqrt{17}-5) c_{2} e^{\sqrt{\frac{51}{2}} r}-(\sqrt{17}+5) c_{1}\right) \tag{B.9}
\end{align*}
$$

where we have suppressed the integration constant of $f$ since it can be absorbed with a rescaling of the coordinates.

Now we can see that we have a superposition of two solutions, one which dominates for $r \gg 0$ and the other one for $r \ll 0$. If we set $c_{1}=0, r \ll 0$ or $c_{2}=0, r \gg 0$, we have that both the solution leads to the IR fixed point. In the first case however the numerics doesn't lead to a well behaving flow, so we consider the second possibility. Notice that if $r$ is big enough we don't need to consider $\epsilon$ very small, since the negative exponential is already a small correction to the IR contribution. For these reason we set $c_{1}=\epsilon=1$ and the initial condition for the plot in figure 1 are given by the linearized expressions at $r=4$.

## B. $2 \quad \mathrm{AdS}_{3}$ flow

The discussion for the $\mathrm{AdS}_{3}$ flow is very similar to the previous one. After the redefinition the BPS system (3.11)-(3.13) reads

$$
\begin{align*}
X^{\prime}+\frac{3}{4} X\left(X^{-2}-X^{2}-\frac{Y}{3}\right) & =0  \tag{B.10}\\
9 X^{2}+3 X^{-2}-\frac{6 Y^{\prime}}{Y}-5 Y & =0  \tag{B.11}\\
4 f^{\prime}+3 X^{2}+X^{-2}+Y & =0 \tag{B.12}
\end{align*}
$$

From (B.10) we can define $Y$ and the system reduces to:

$$
\begin{equation*}
f^{\prime}=-\frac{1}{X^{2}}-\frac{X^{\prime}}{X}, \quad 6 X^{\prime \prime}+\frac{14 X^{\prime 2}}{X}+\left(\frac{18}{X^{2}}-48 X^{2}\right) X^{\prime}+18 X^{5}-27 X+\frac{9}{X^{3}}=0 \tag{B.13}
\end{equation*}
$$

We can now solve this equation linearizing it close to the fixed point:

$$
\begin{equation*}
X=\sqrt[4]{\frac{1}{2}}+\epsilon x(r) \tag{B.14}
\end{equation*}
$$

At the first order in $\epsilon, x$ has to solve

$$
\begin{equation*}
x^{\prime \prime}-\sqrt{2} x^{\prime}-6 x=0 \tag{B.15}
\end{equation*}
$$

which admits the solution

$$
\begin{equation*}
x=e^{-\frac{(\sqrt{13}-1) r}{\sqrt{2}}}\left(c_{2} e^{\sqrt{26} r}+c_{1}\right) \tag{B.16}
\end{equation*}
$$

Using this definition of $X$ we can derive the ones of $f$ and $Y$ at the first order in $\epsilon$ :

$$
\begin{align*}
& Y=\frac{3}{\sqrt{2}}+\epsilon 2^{3 / 4} e^{-\sqrt{7-\sqrt{13}} r}\left((2 \sqrt{13}-7) c_{2} e^{\sqrt{26} r}-(2 \sqrt{13}+7) c_{1}\right)  \tag{B.17}\\
& f=-\sqrt{2} r+\epsilon \frac{1}{3} \sqrt[4]{2} e^{-\frac{(\sqrt{13}-1) r}{\sqrt{2}}}\left((\sqrt{13}-4) c_{2} e^{\sqrt{26} r}-(\sqrt{13}+4) c_{1}\right) \tag{B.18}
\end{align*}
$$

Now we have again two solutions. The IR fixed point is at $c_{2}=0, r \gg 0$. The initial point for the numerical analysis in figure 2 is given by the linear expressions at $c_{1}=\epsilon=1$ and $r=5$.

## B. $3 \quad \mathrm{AdS}_{2}$ flow

In this case we adopt the definition same definitions of $g$ and $\tilde{g}$ as before, moreover, as discussed in section 3.3, we set $h_{1}=h_{2}=-\frac{1}{2} \log Y$. The BPS system therefore reduces to a set of three equations

$$
\begin{align*}
X^{\prime}+\frac{3}{4} X\left(X^{-2}-X^{2}\left(1-\frac{Y^{2}}{27}\right)-\frac{2 Y}{9}\right) & =0  \tag{B.19}\\
\left(9-\frac{Y^{2}}{3}\right) X^{2}+3 X^{-2}-\frac{6 Y^{\prime}}{Y}-2 Y & =0  \tag{B.20}\\
12 f^{\prime}+\left(9+Y^{2}\right) X^{2}+3 X^{-2}+2 Y & =0 \tag{B.21}
\end{align*}
$$

We can use (B.19) to define $Y$, there are two possible solutions, but the one which leads to the $\mathrm{AdS}_{6}$ fixed point is

$$
\begin{equation*}
Y=\frac{3}{X^{2}}\left(1-\sqrt{3 X^{4}-4 X X^{\prime}-2}\right) . \tag{B.22}
\end{equation*}
$$

Using this definition we have

$$
\begin{align*}
& f^{\prime}=\frac{3 X X^{\prime}+2 \sqrt{3 X^{4}-4 X X^{\prime}-2}-3 X^{4}}{X^{2}}  \tag{B.23}\\
& X^{\prime \prime}-7 \frac{X^{\prime 2}}{X}+3\left(X^{4}-2 \sqrt{3 X^{4}-4 X X^{\prime}-2}\right) \frac{X^{\prime}}{X^{2}}+\frac{3 X^{4}-2}{X^{3}}\left(\sqrt{3 X^{4}-4 X X^{\prime}-2}-1\right)=0 \tag{B.24}
\end{align*}
$$

Again, we linearize this equations near to the fixed point solution

$$
\begin{equation*}
X=\sqrt[4]{\frac{1}{2}}+\epsilon x(r) \tag{B.25}
\end{equation*}
$$

where

$$
\begin{equation*}
x=c_{1} e^{-2 \sqrt{6} r}+c_{2} e^{\sqrt{6} r} . \tag{B.26}
\end{equation*}
$$

From this definition it automatically follows that

$$
\begin{equation*}
Y=3 \sqrt{\frac{3}{2}}-\sqrt[8]{3^{17} 2^{7}} \sqrt{c_{1} \epsilon} e^{-\sqrt{6} r}, \quad f=-\sqrt{6} r-\sqrt[8]{3^{5} 2^{11}} \sqrt{c_{1} \epsilon} e^{-\sqrt{6} r} \tag{B.27}
\end{equation*}
$$

The IR fixed point is now obtained at $c_{2}=0, r \gg 0$. The initialization parameters for figure 3 are given by the expressions above at $c_{1}=\epsilon=1$ and $r=5$.

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[^0]:    ${ }^{1}$ Choosing $\tilde{g}=3 / \sqrt{2}$ gives the metric on $\mathrm{AdS}_{6}$ with a unit radius. We can choose to work with a generic $\tilde{g}$ by adding a constant factor in the expansion of $f$ modifying the $\operatorname{AdS}_{6}$ radius.

[^1]:    ${ }^{2}$ A lift to Massive IIA was presented in [63].

[^2]:    ${ }^{3}$ In the balanced case they must satisfy the relation $F_{i}=2 N_{i}-N_{i-1}-N_{i+1}$.

